# Light (anti)Nucleus Production in $\sqrt{s_{NN}} = 7.7 - 200 \text{ GeV}$ Au+Au Collisions in the STAR Experiment

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Light (anti)nucleus production in 7.7 – 200 GeV Au+Au collisions in STAR Introduction



A complete understanding of the hot and dense partonic and/or nuclear systems formed at RHIC requires an understanding of the latest stages of the collisions.

Light nucleus rates and spectra are a "direct" nucleon correlation observable that is complementary to two-particle correlations obtained, *e.g.*, from intensity interferometry (HBT).

Thermodynamic approaches and the sudden approximation can provide insight on source "homogeneity volumes" and emission profiles, (anti)proton phase space densities, *etc*.



#### Cuts

Outlier run rejection based on multiple global observables.

$\sqrt{s_{_{NN}}}$	Run	N <sub>events</sub>
7.7	2010	5M
11.5	2010	15M
19.6	2011	37M
27	2011	46M*
39	2010	58M*
62.4	2010	59M*
200	2010	51M*
200	2011	47M*

\*Not entirety of available data.

**Event Cuts**  $|Z_{vtx}| < 50$  cm for  $\sqrt{s_{NN}} \le 39$  GeV,  $|Z_{vtx}| < 30$  cm otherwise  $R_{vtx} < 2cm$ Pileup event rejection based on multiple global observables.

## **Primary Track Cuts**

 $N_{hitsfit} > 15$  (of 45 possible)  $N_{hitsdedx} > 10$  (of ~35 possible) Global partner D.C.A. to primary vertex < 3 cm TOF: "good match" criterion  $\geq 1$ TOF:  $Y_{local} < 1.8$  cm

## Centrality

Uses primary track multiplicity within  $|\eta| < 0.5$ Corrected for  $Z_{vtx}$  and beam luminosity dependence

### **Particle Identification**

Uses TPC dE/dx and Time Of Flight (TOF) ...Careful avoidance of dE/dx "merged tracks" Statistical, in small (P<sub>T</sub>,y,centrality) bins

Uncertainties are statistical only.





Efficiencies depend on year,  $\sqrt{s_{NN}}$ , centrality, species, rapidity, P<sub>T</sub>







T. F. Hoang, *et al.*, Z. Phys. C29, 611 (1985) Check material budget via p & pbar embedding

(anti)proton Feed-down  $(\Lambda, \Sigma \rightarrow p)$ UrQMD 3.3p1 simulations full reconstruction





Sharp increase in nucleus cross-sections at low  $M_T$  is due to spallation:

X + Beam Pipe = p,d,t + Y

Significant for  $P_T < \sim 0.5 * A$ , does not produce antinuclei

Light (anti)nucleus production in 7.7 – 200 GeV Au+Au collisions in STAR Coalescence Ratio B<sub>2</sub>

 $B_2 = \sigma_d / \sigma_p^2$ , where the cross-sections are evaluated at the same velocity (P<sub>T</sub>/A)

 $B_2$  is a dimensioned ratio that can be related in model-dependent ways to a "homogeneity volume":  $B_A \sim 1/V$ 

WJL, S. Pratt et al., Phys. Rev. C 52, 2004 (1995), R. Scheibl & U. Heinz, Phys. Rev. C 59, 1585 (1997)



Lines are UrQMD 3.3p1 or Pythia model calculations plus a "dynamic coalescence afterburner" uses 6D coalescence with one of three d wave functions for A+A (UrQMD), 3D coalescence for Pythia. J. L Nagle et al., Phys. Rev. C **53**, 367 (1996), B. Monreal, WJL, *et al.*, Phys. Rev. C **60**, 31901 (1999)





Proton density falls with  $\sqrt{s_{NN}}$ , Antiproton density rises with  $\sqrt{s_{NN}}$ Trends reflect decreasing baryochemical potential,  $\mu_B$ , with increasing  $\sqrt{s_{NN}}$ 



Light (anti)nucleus production in 7.7 –

Source "Homogeneity" Radii

Light (anti)nucleus production in 7.7 – 200 GeV Au+Au collisions in STAR

$$B_2 = \frac{\sigma_d}{\sigma_p^2}, \quad (R_G^2 + \frac{\delta^2}{2})^{3/2} = \frac{3}{2} \frac{\pi^{3/2} \hbar^3}{B_2 m_p c^2} \quad \text{(anti)deuteron Gaussian width: } \delta \sim 2 \text{ fm}$$
$$m_p = \text{proton mass}$$



HBT results from H. Zbroszczyk, 7<sup>th</sup> WPCF, 2011 http://tkynt2.phys.s.u-tokyo.ac.jp/wpcf2011/talks/sept20/Zbroszczyk.pdf



- Light (anti)nuclei have been measured in Au+Au collisions at seven beam energies by STAR at RHIC.
- Spectra versus  $P_T$ ,  $P_T/A$ ,  $M_T/A$ , and  $M_T-M_0$  provide information on the nucleon source near freeze-out.
- Hardening of the spectra with the mass number reflects strong transverse flow.
- Qualitative reproduction of B<sub>2</sub> values by UrQMD+dynamic coalescence calculation.
- Gaussian radii from B<sub>2</sub> values similar to that from (anti)proton intensity interferometry (HBT).
- Antiproton and proton phase space densities approach each other as  $\sqrt{s_{NN}}$  increases, reflecting decreasing  $\mu_B$ .

