



# Modern Time-Of-Flight

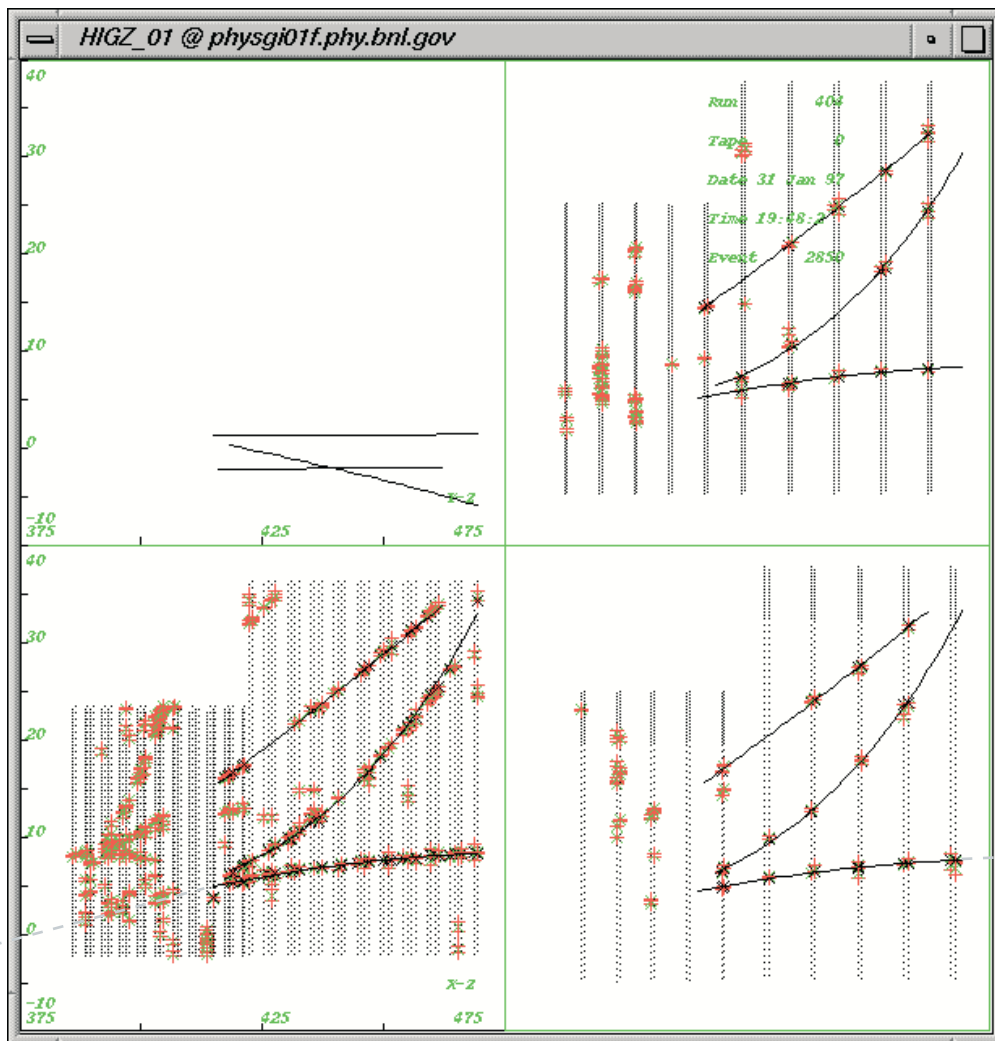
W.J. Llope  
*Rice University*

*Cyclotron Institute Colloquium  
TAMU, College Station  
March 5, 2002*

## OUTLINE:

- TOF philosophy & Particle IDentification...
- Examples of Conventional TOF detectors  
E896 TOF and STAR TOFp/pVPD
- Other PID techniques, & physics topics with  
no TOF, a little TOF, and a lot of TOF.
- The looming problem:  
general need for large area TOF, but  
large area Conventional TOF can be *expensive*
- A new technology → towards “Modern” TOF  
...as fast, easy to build, and <1/10th the price...
- New production prototype for STAR: “TOFr”
- Summary and outlook...  
...towards STAR TOFR (tens of sq. meters)...

“Tracking” charged particles through a magnetic field....



Information available following “reconstruction” of a track

- track **trajectory**....
- track **momentum** components...
- (w/ some “extrapolation”) track total **path length** between two points

One can furthermore directly identify each track using Time-Of-Flight  
 → new or supplemental capabilities for Particle IDentification (PID)

To assign a mass for a given track, measure a **time** at one point along this track relative to a specific reference time.

**time** = stop time - start time (and lots of corrections)

**distance** = rate \* **time**

→  $\beta = S/\Delta t/c$

**momentum** = mass \* rate

→  $m = p/(\gamma\beta)$

## Time of Flight Reason for Being: (charged) Particle Identification

Things you can do with a TOF system (to be illustrated on the coming slides)

- Optimization of PID capabilities of other detectors for tracks in a common acceptance
- Spectra for primary tracks: transverse momentum and rapidity distributions by species
- Searches: significant suppression to backgrounds by requiring the 2(3) tracks assumed to be daughters of a decaying particle of interest are each the correct species for this assumption.

Depending on what you're trying to identify, you'd need....

Sufficiently accurate **tracking**

$\sigma(p)/p \sim$  few percent,  $\sigma(S) \sim$  few mm-cm

Sufficiently long **flight paths**

$S \sim$  few meters

**Start detector** with sufficient resolution

$\sigma(t_{\text{start}}) \sim 20\text{-}60\text{ps}$

**Stop detector** with sufficient resolution

$\sigma(t_{\text{stop}}) \sim 50\text{-}100\text{ps}$

**Digitization** with sufficient resolution

$\sim 100\text{ns}/\sim 11\text{bits}$ ,  $\sigma \sim 20\text{ps}$

Lots of **software** for the corrections

Slewing, Timing offsets, conversion factors, ...

$\sigma(\text{corrections}) \sim 20\text{-}30\text{ps}$

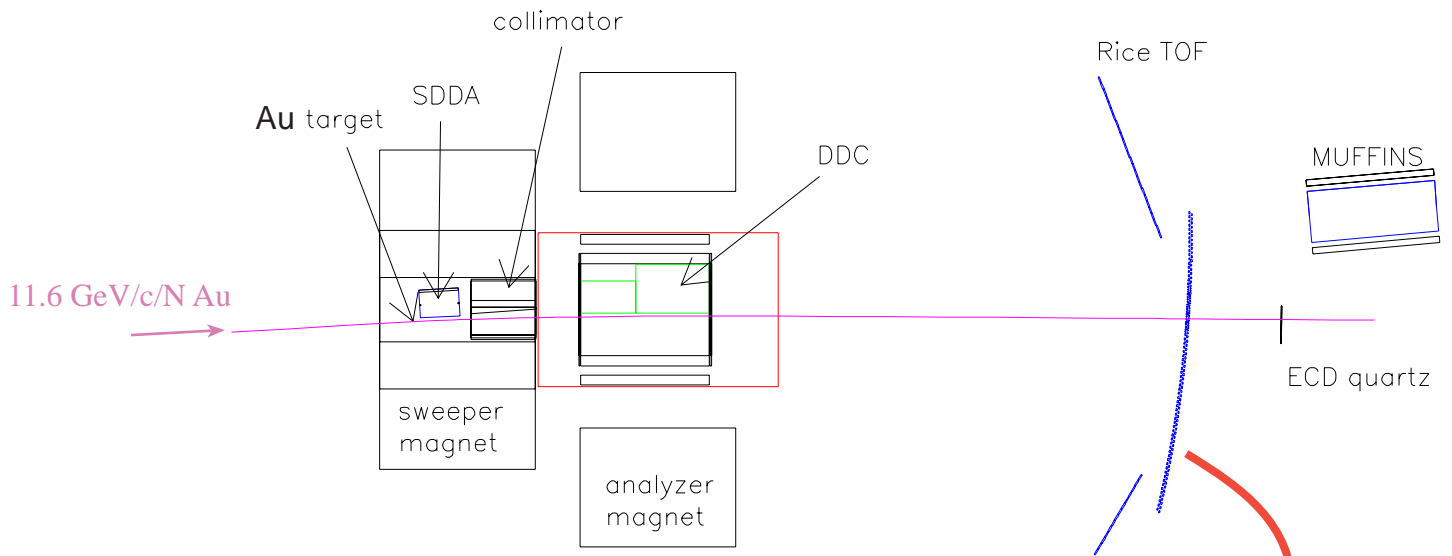
**Monitoring**: rate scalars, temperature along cable paths, etc. etc.....

Total resolution on time differences is typically 70-150ps after all the corrections

time differences between different particle species (pion, Kaon, proton, fragments) can be large!

Now to a few concrete examples.....

# “Conventional” TOF Example 1: BNL-AGS E896 TOF



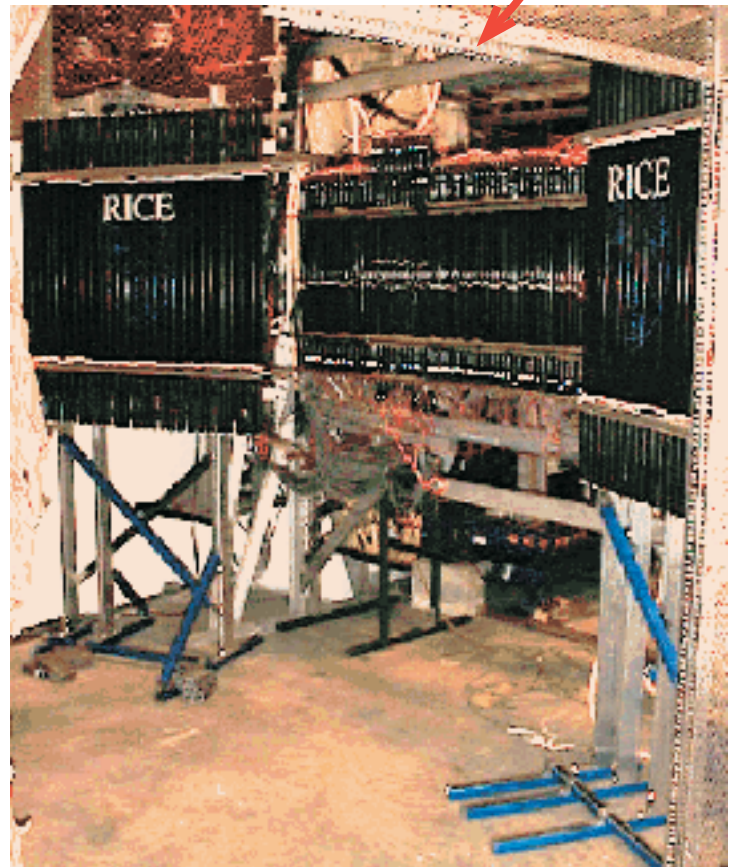
Search for the H,  $\Lambda$ , Ks using topology and momenta only

**BUT....** no direct PID on the charged daughters

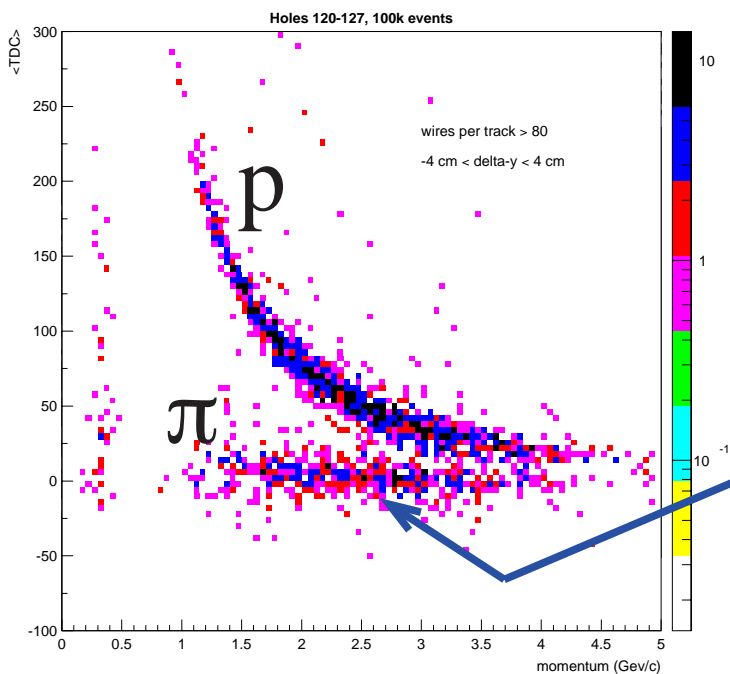
does each candidate  $\Lambda \rightarrow p \ \& \ \pi^-$ ??

does each candidate  $Ks \rightarrow \pi^+ \ \& \ \pi^-$ ??

does each candidate  $H \rightarrow p \ \& \ X$ ??  
( $X = \Sigma^- \rightarrow n \ \& \ \pi^-$ )



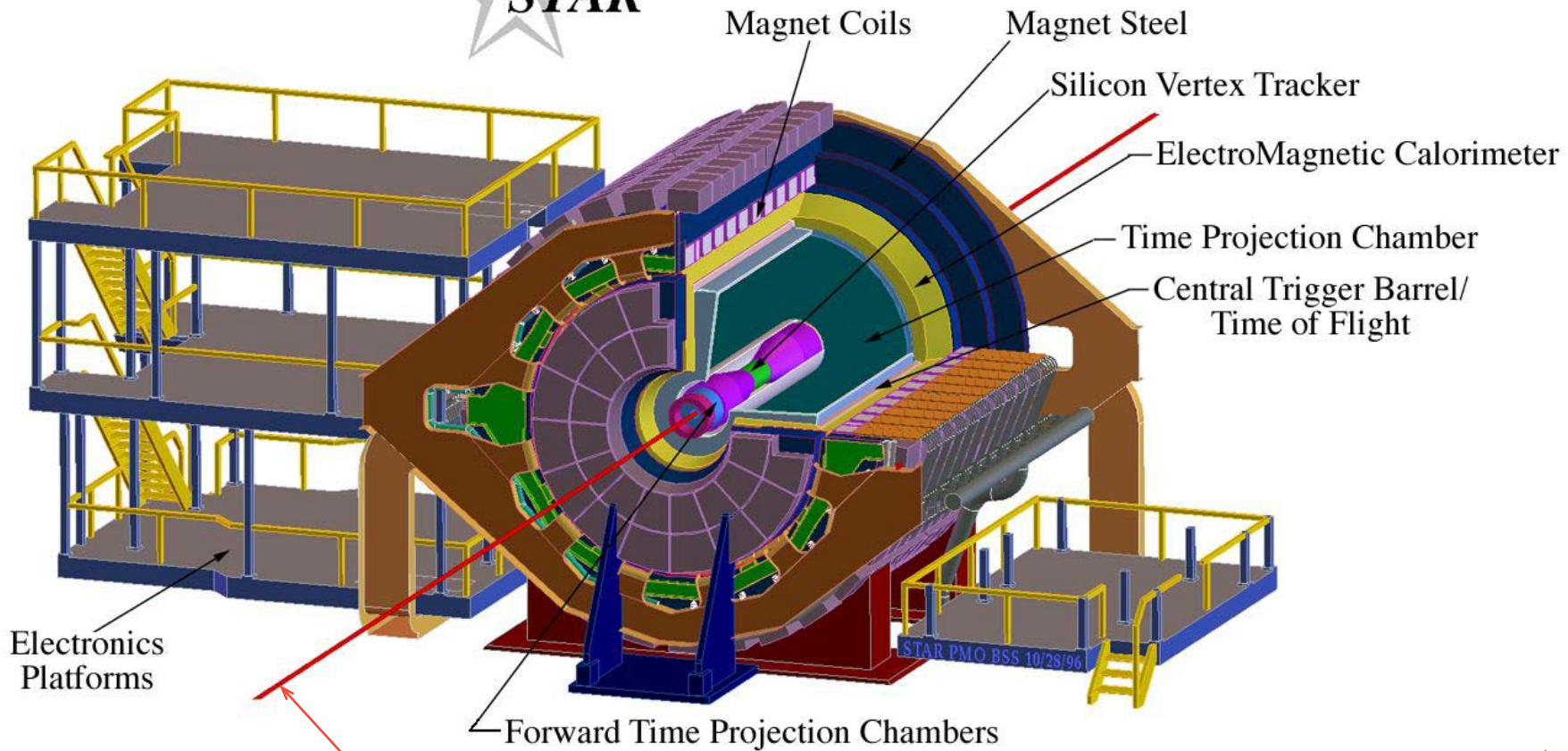
E896 candidates:  $\Lambda?$   $\rightarrow$  positive....



- ~10% of E896's best  $\Lambda$  candidates are *not*  $\Lambda$ 's...

1/3 are really Ks  
2/3 are other backgrounds

- $\Lambda$  and Ks spectra w/ daughter PID  
(K. Kainz, Ph.D. Thesis, 2001)



Electronics  
Platforms

Forward Time Projection Chambers

ALICE at CERN LHC

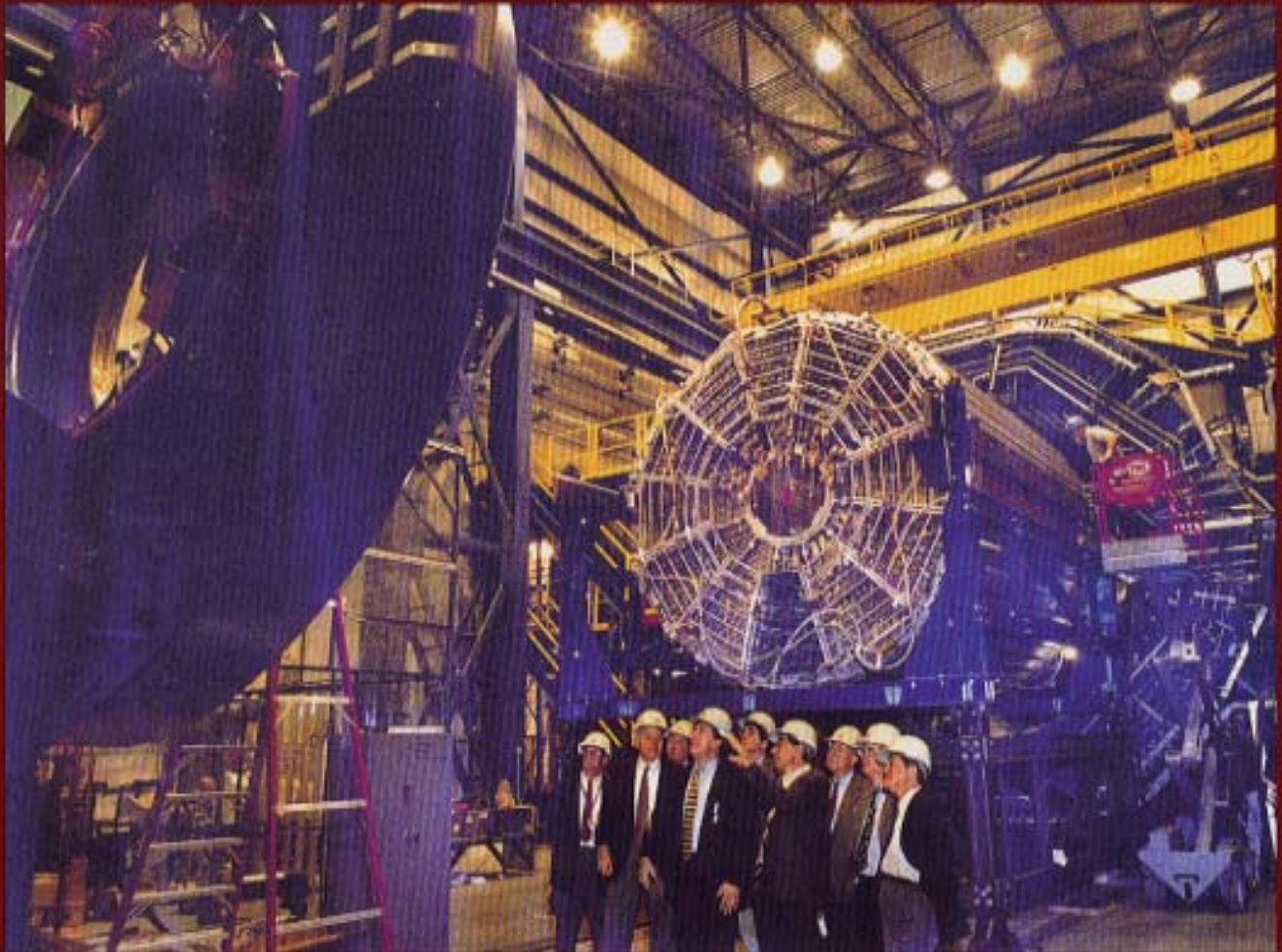
100 GeV/c/N Au + 100 GeV/c/N Au  
10 GeV/c/N Au + 10 GeV/c/N Au  
100 GeV/c p + 100 GeV/c p  
and there will be others!



INTERNATIONAL JOURNAL OF HIGH-ENERGY PHYSICS

# CERN COURIER

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## A STAR in the making

### LHC COLD START

Procurement and construction work for CERN's LHC project begin in earnest

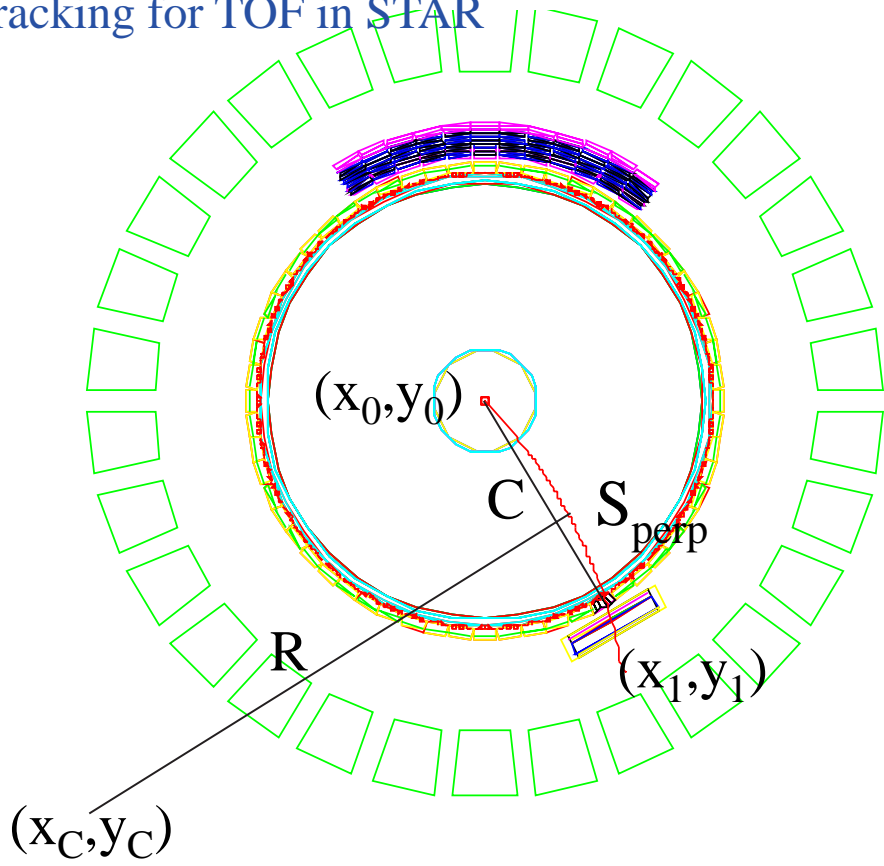
### EPIC DEVELOPMENTS

Physics helps and is helped by a new generation of microprocessors being tested at CERN

### MYSTERIES OF COSMIC RAYS

Unexplained phenomena in the ultra-high-energy area point to gaps in our understanding

# Tracking for TOF in STAR



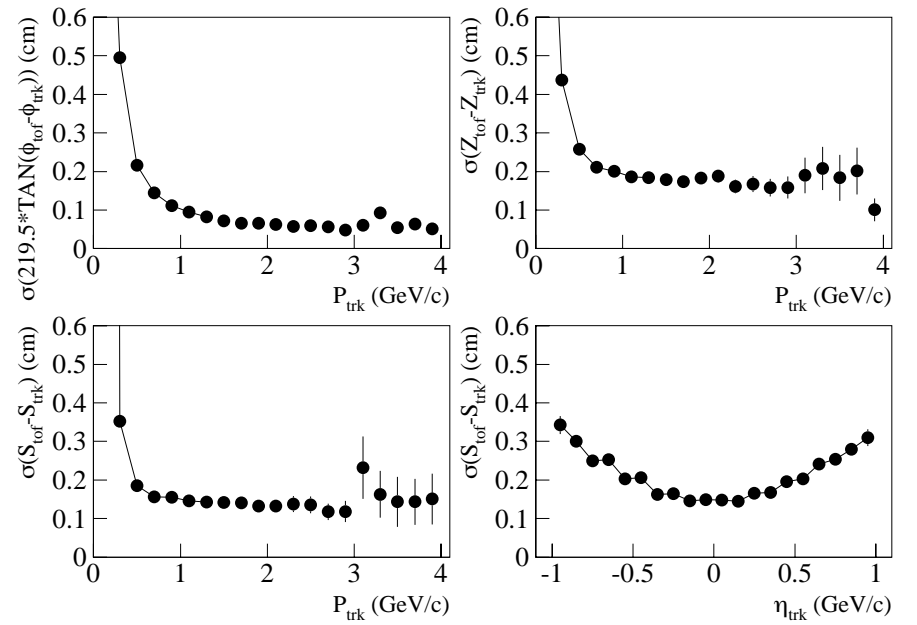
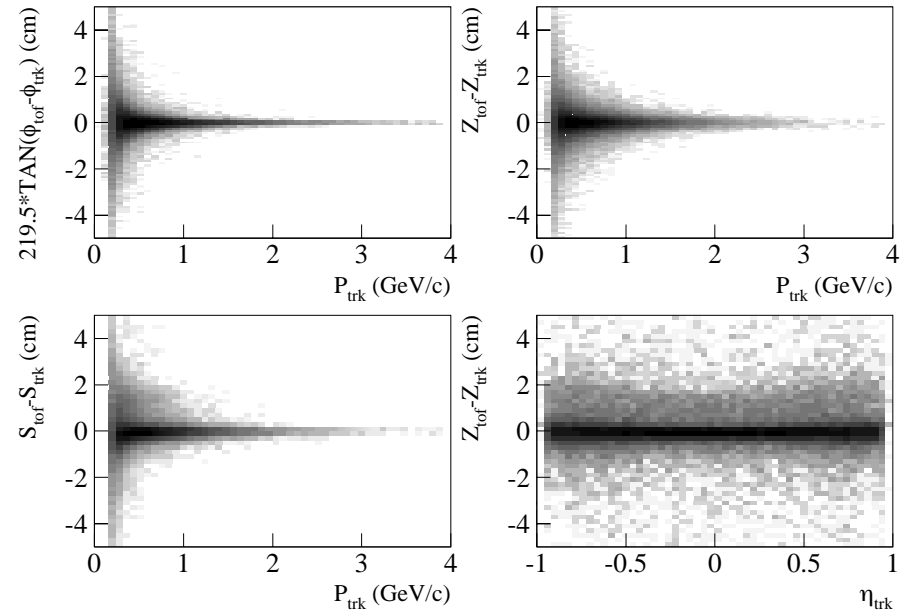
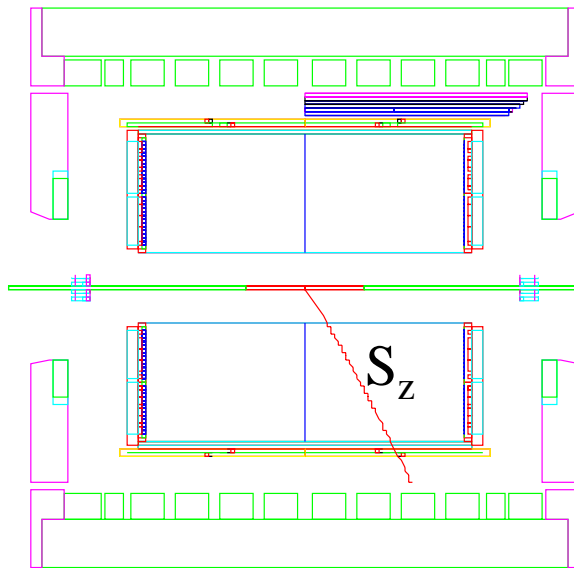
$$s = \sqrt{(S_{\text{perp}})^2 + s_z^2} \text{ with}$$

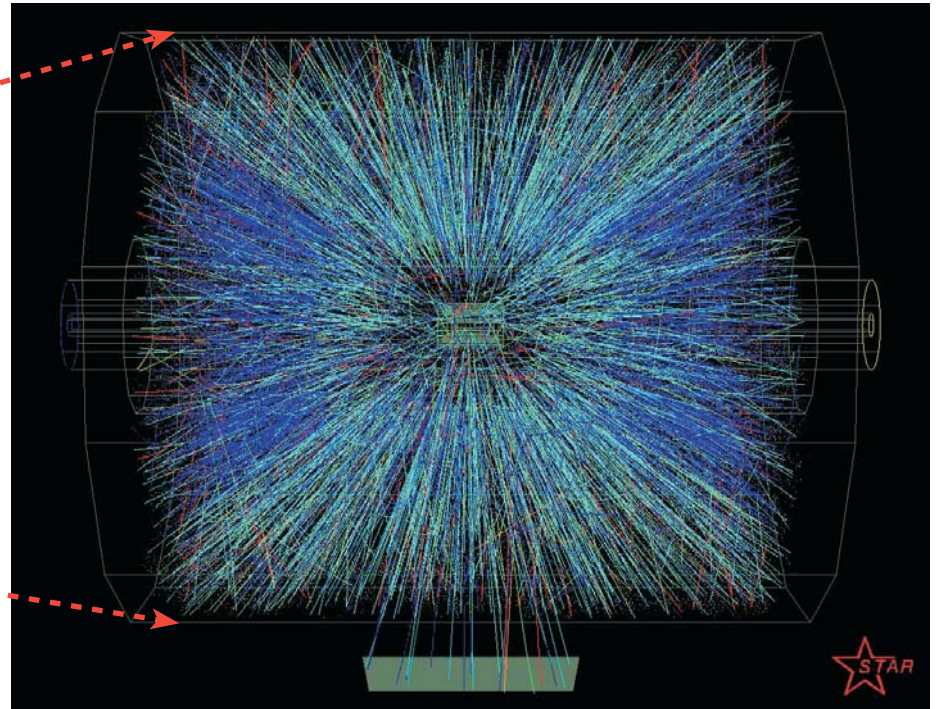
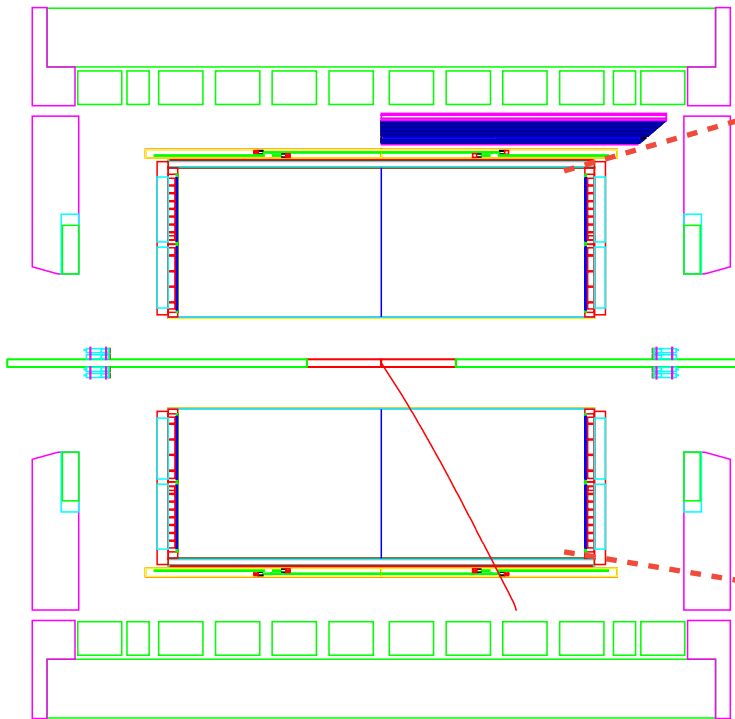
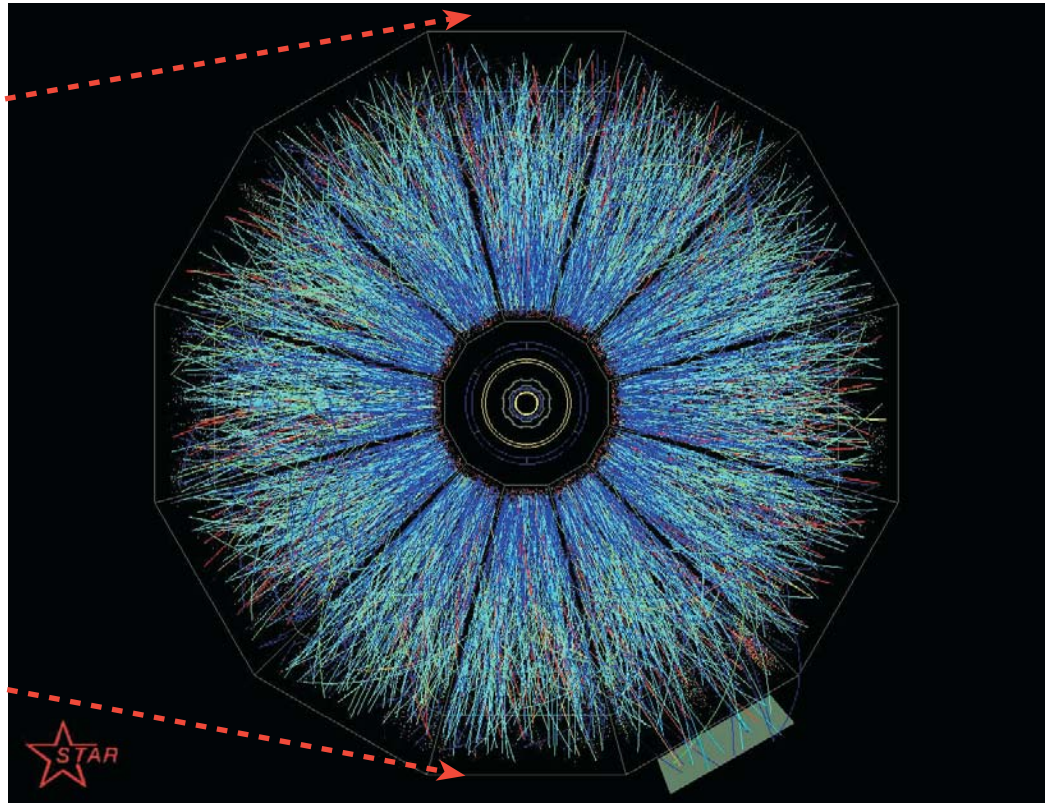
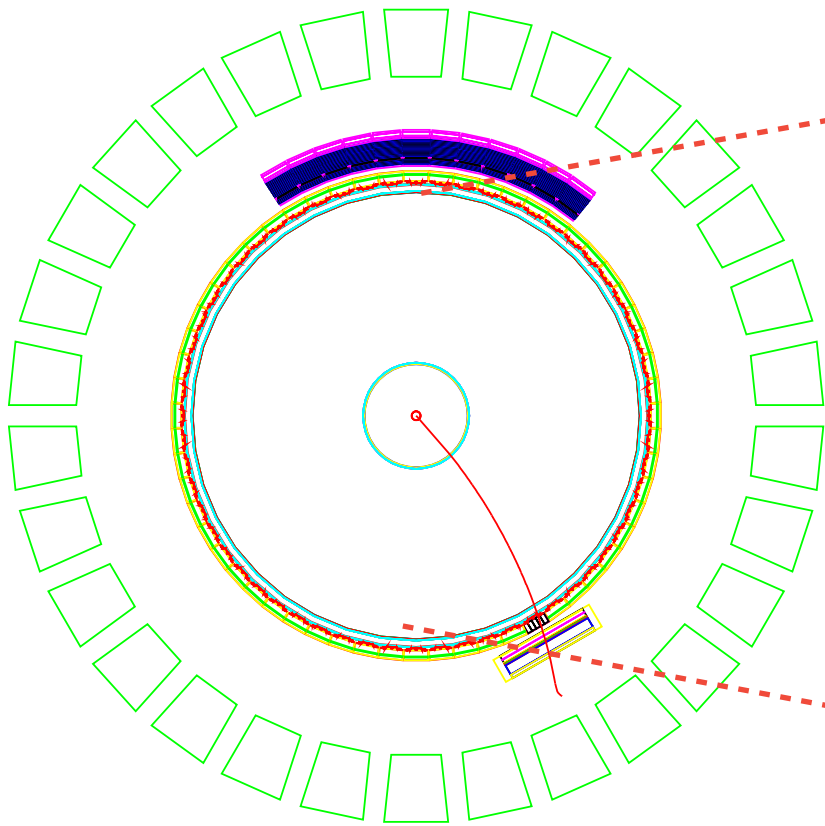
$$S_{\text{perp}} = 2R \sin^{-1}(C/2R)$$

$$s_z = Z_1 - Z_0$$

$R$  is the helix radius

$$C = \sqrt{((x_1 - x_0)^2 + (y_1 - y_0)^2)}$$

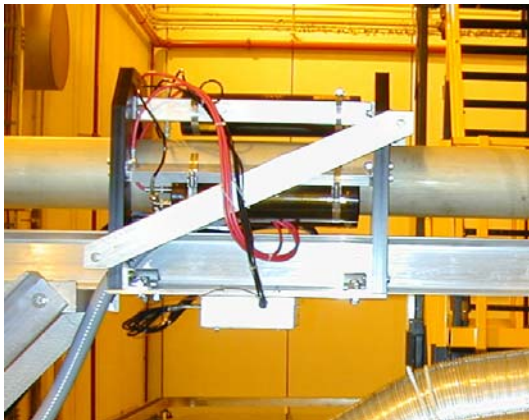




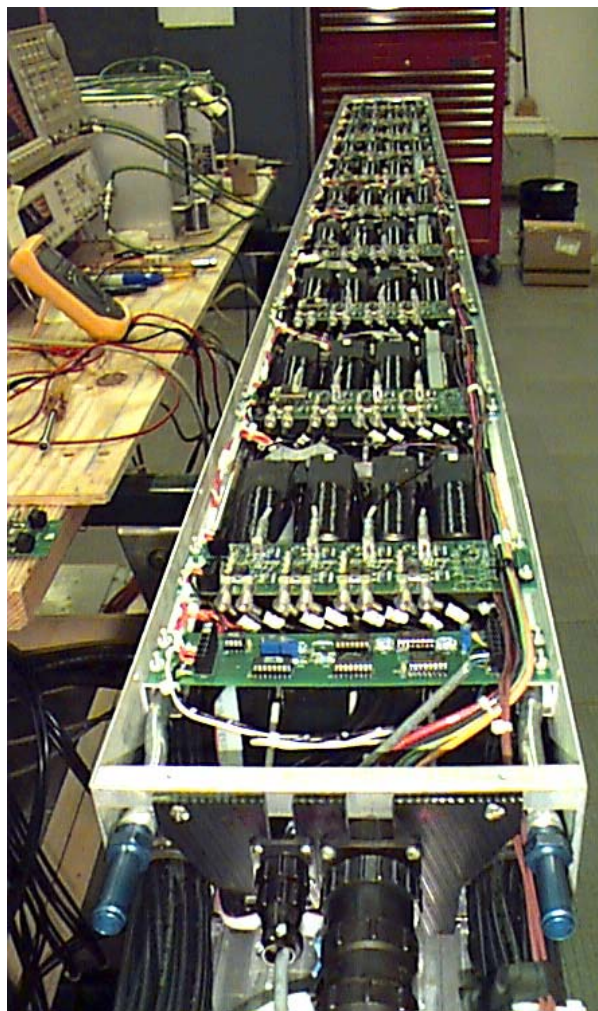


## Conventional TOF Example 2: STAR TOFp

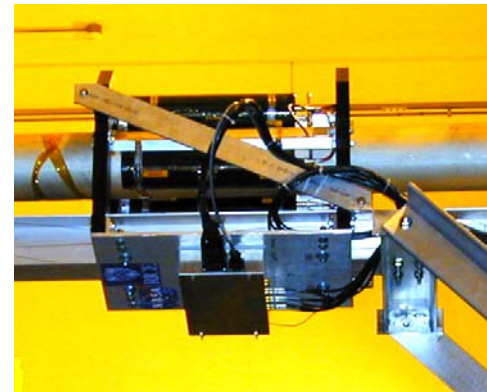
*pVPD East*



*TOFp tray*



*pVPD West*



### Hardware

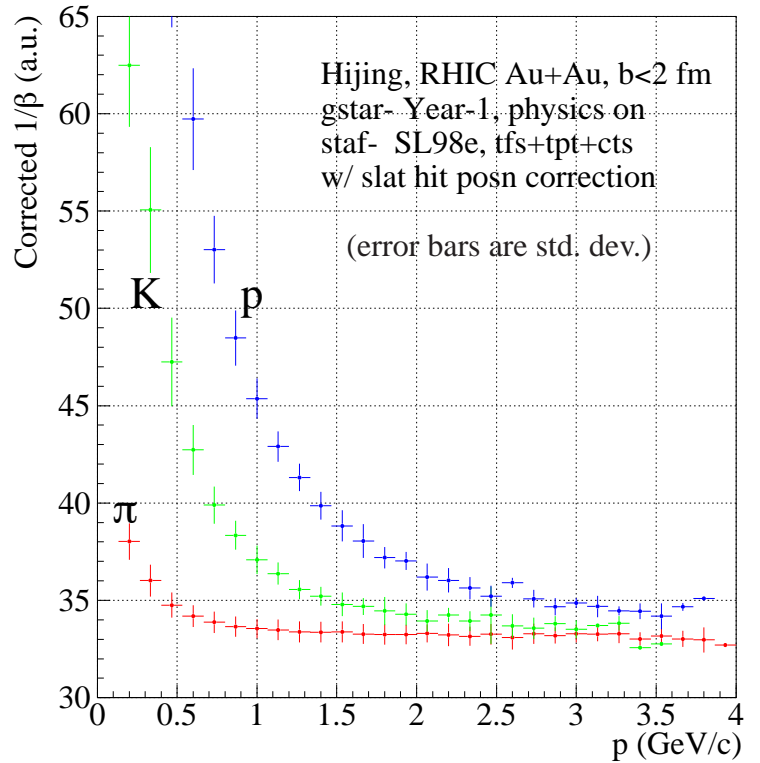
- Start: 6 pVPD channels
- Stop: 41 TOFp channels
- Local trigger: scalers
- Temp. monitoring: 28 TCs
- Correlated noise: 60Hz ramp

### Operations

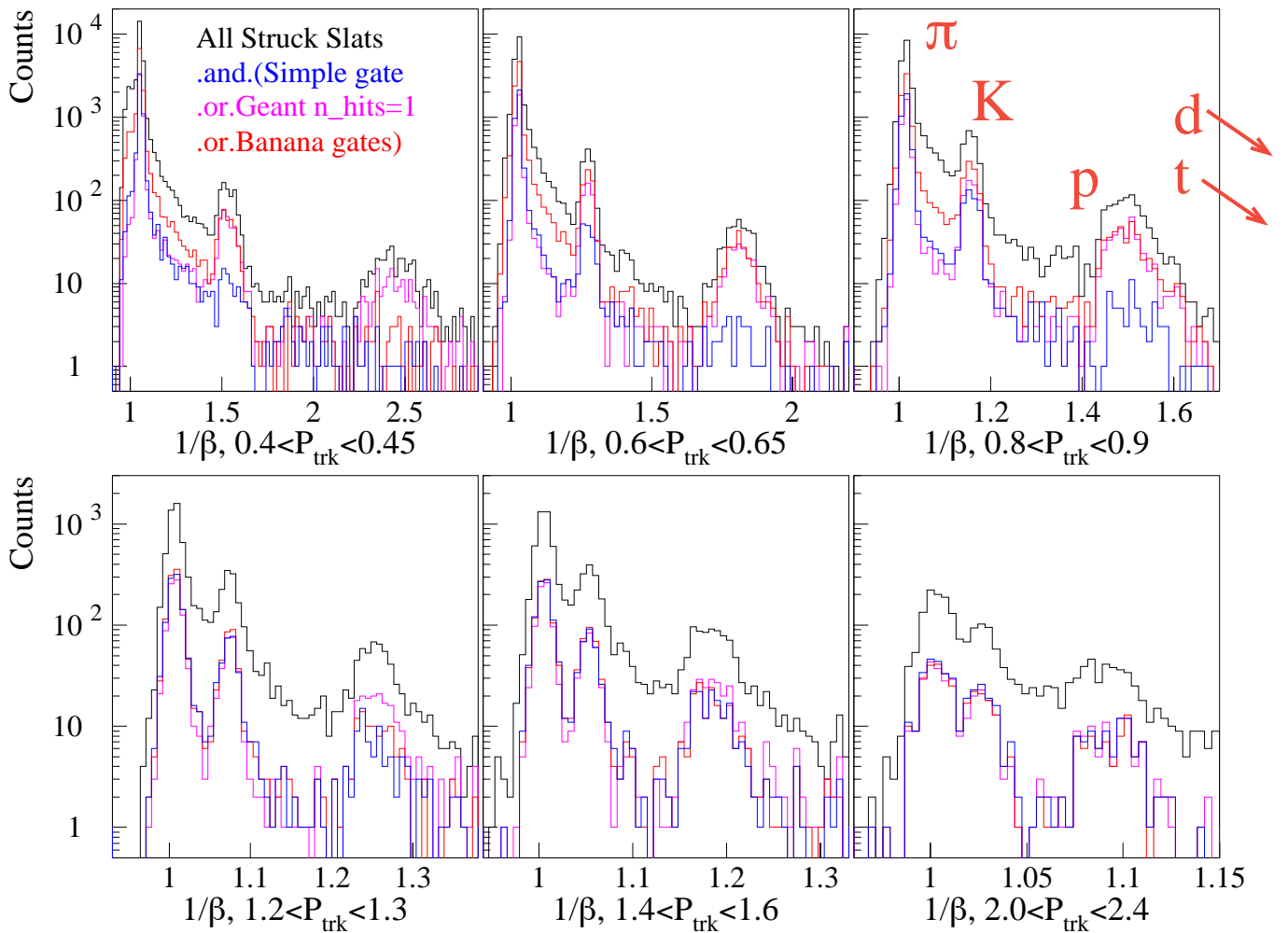
- pVPD active since 8/25
- TOFp active since 10/9
- Collected events:
  - AuAu Central (200GeV): 2.7M
  - AuAu Minbias(200GeV): 1.5M
  - AuAu Central (22GeV): 72k
  - AuAu Minbias (22GeV): 216k
  - pp (200GeV): 21.8M (100.00%)

Given STAR's geometry  
& a 100ps TOF system:

- $\pi/K/p$  direct PID for  $\sim 0.3 \text{ GeV}/c < p < 1.7-1.9 \text{ GeV}/c$
- $(\pi+K)/p$  direct PID for  $\sim 0.3 \text{ GeV}/c < p < 2.8-3.0 \text{ GeV}/c$

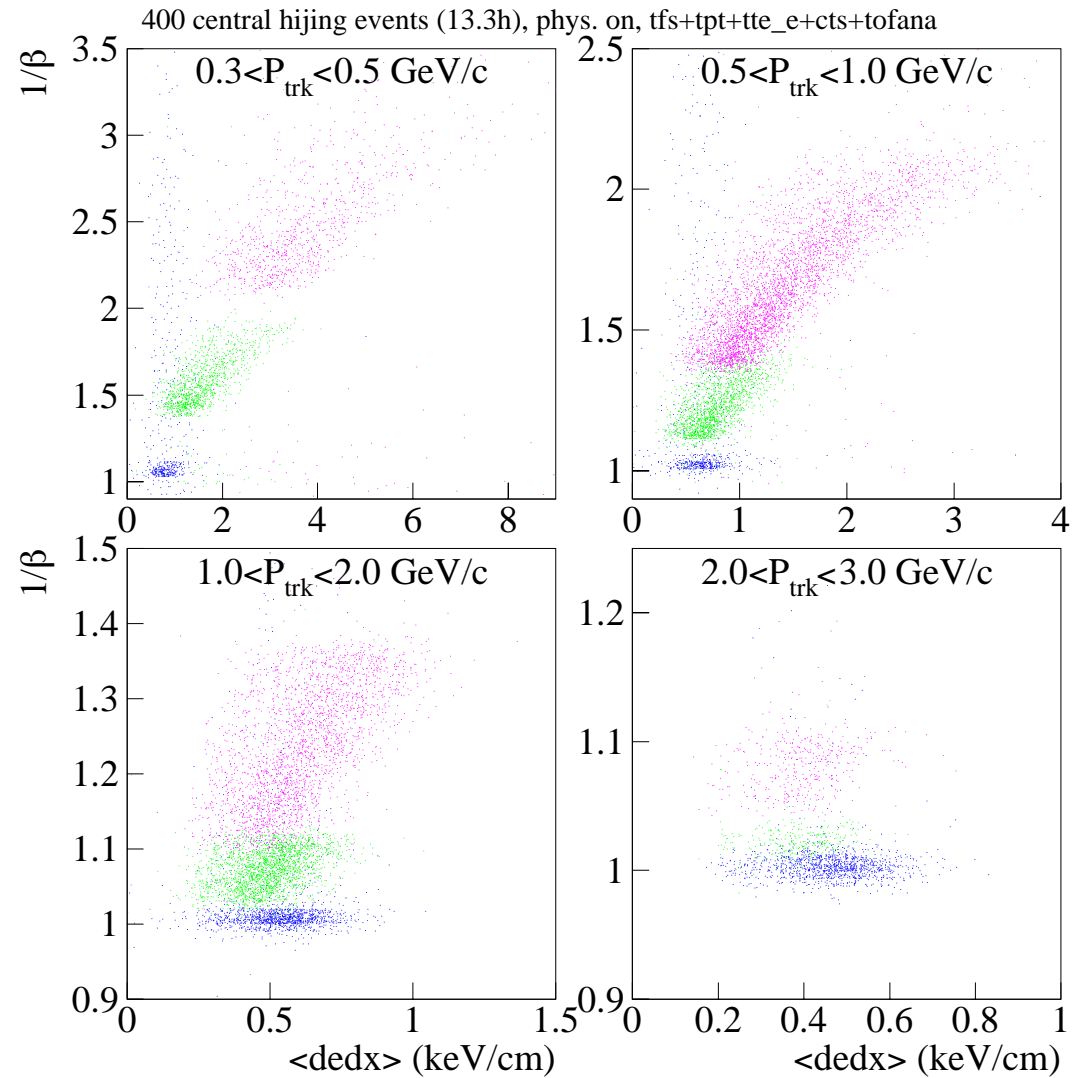
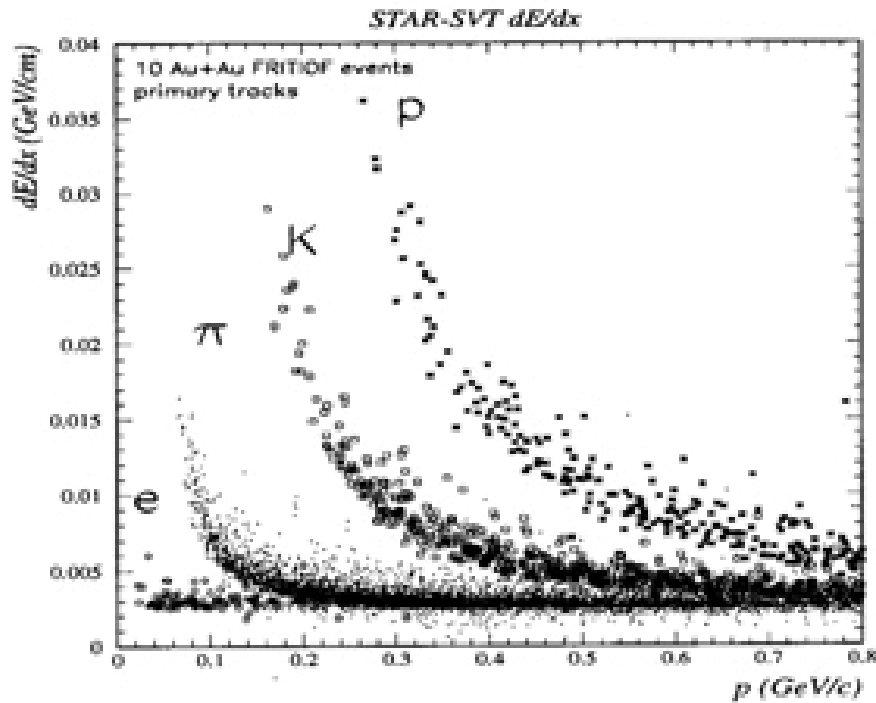


Venus,  $b < 2 \text{ fm}$ , all phys on, 13.3hrs at 1Hz w/ 1 tray

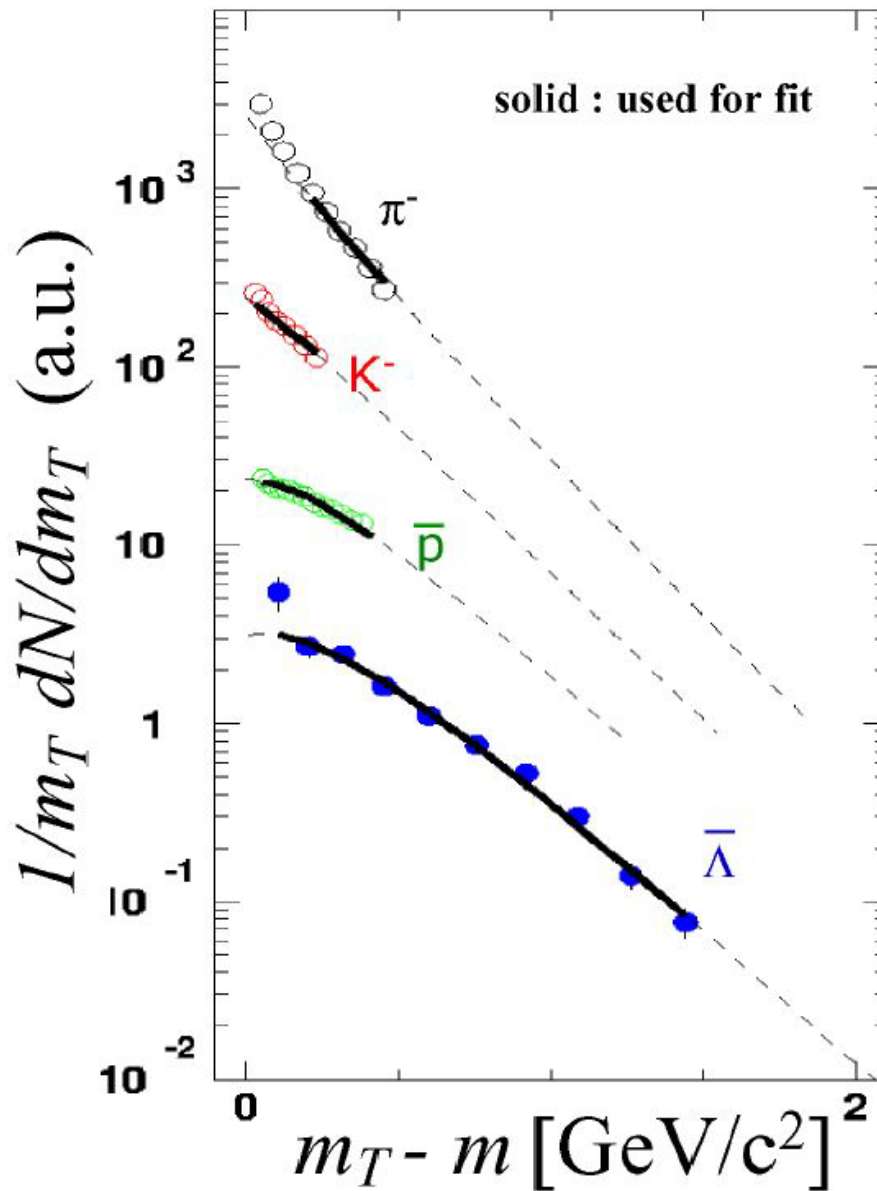


## PID using other kinds of detectors: $dE/dx$ .....

- $\pi/K/p$  direct PID for  
 $\sim 0.1 \text{ GeV}/c < p < 0.6 \text{ GeV}/c$
- $(\pi+K)/p$  direct PID for  
 $\sim 0.1 \text{ GeV}/c < p < 1.0 \text{ GeV}/c$



example of STAR's reach w/ TPC dE/dx alone...



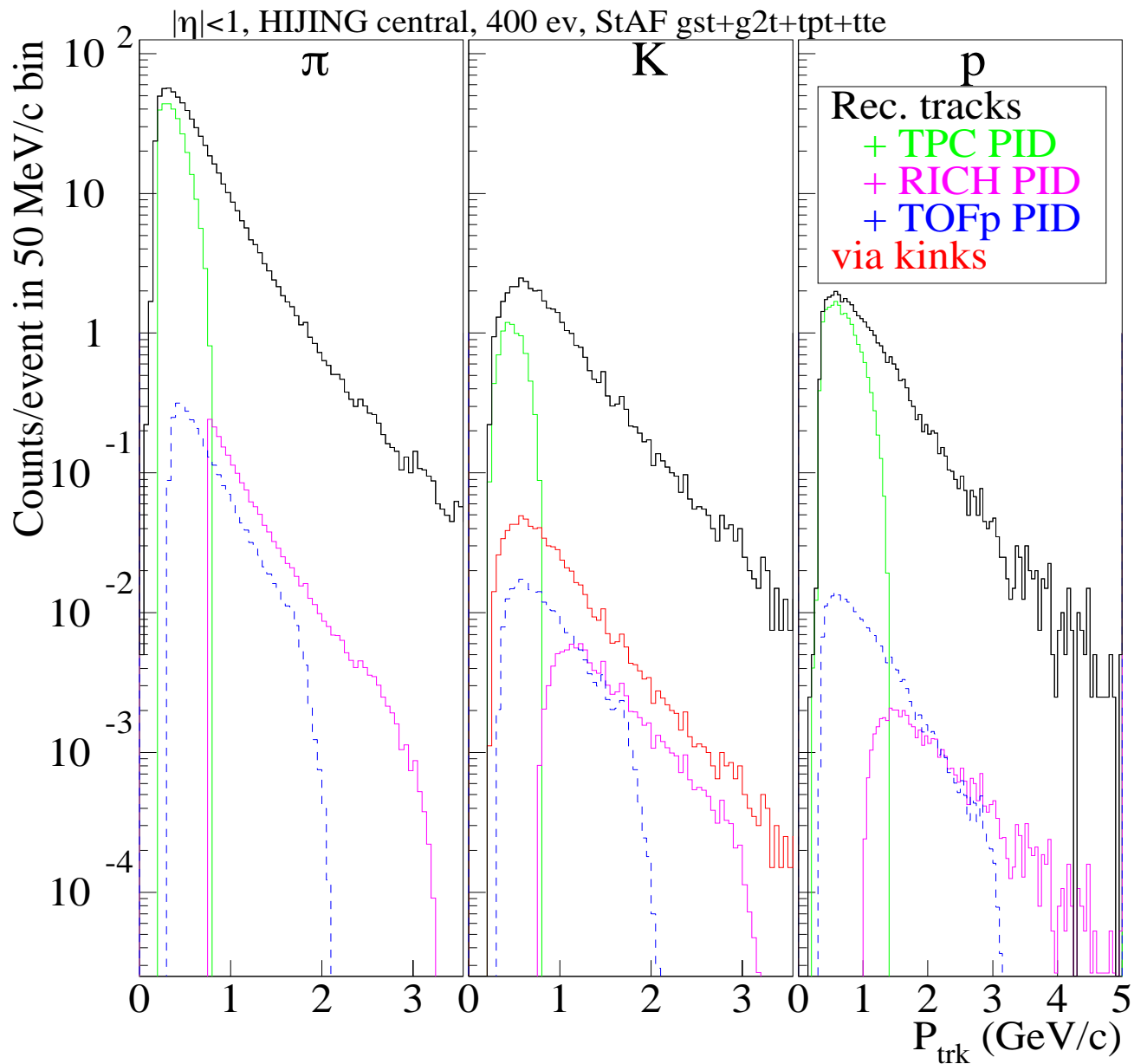
plot by  
M. Kaneta

- acceptance w/ direct PID is in a region of momentum that is *soft*...
- uncomfortably *long* extrapolation to higher momenta...

TOFp (conventional & small area) should extend the reach of this plot by a factor of 2.

...but a small area TOF such as TOFp only gets you so far...

## PID using other kinds of detectors (cont.)



Examples from STAR (for full energy Au+Au collisions)

TPC  $dE/dx$  directly identifies  $\sim 60\%$  of tracks it can reconstruct  
w/ a  $\sim 100\text{ps}$  TOF system in STAR's geometry,

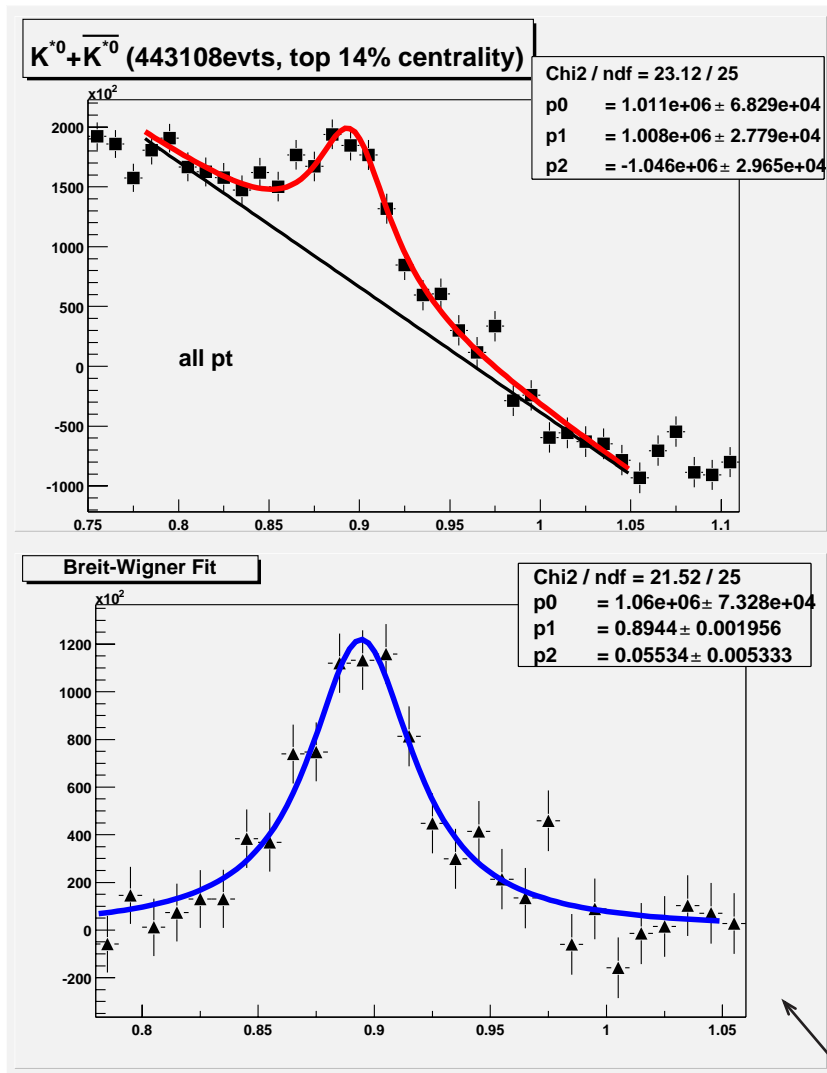
$\rightarrow \sim 97\%$  of reconstructed tracks can be directly identified

want charged hadron PID at even higher momenta?

$\rightarrow$  you want a RICH detector... (see *e.g.* M. Murray)

# Physics requiring Large-Area TOF in STAR, example 1

$K^{*0}(892)$  production in Relativistic Heavy Ion Collisions at  $\sqrt{s_{NN}} = 130 \text{ GeV}$



(plots by Zhangbu Xu)

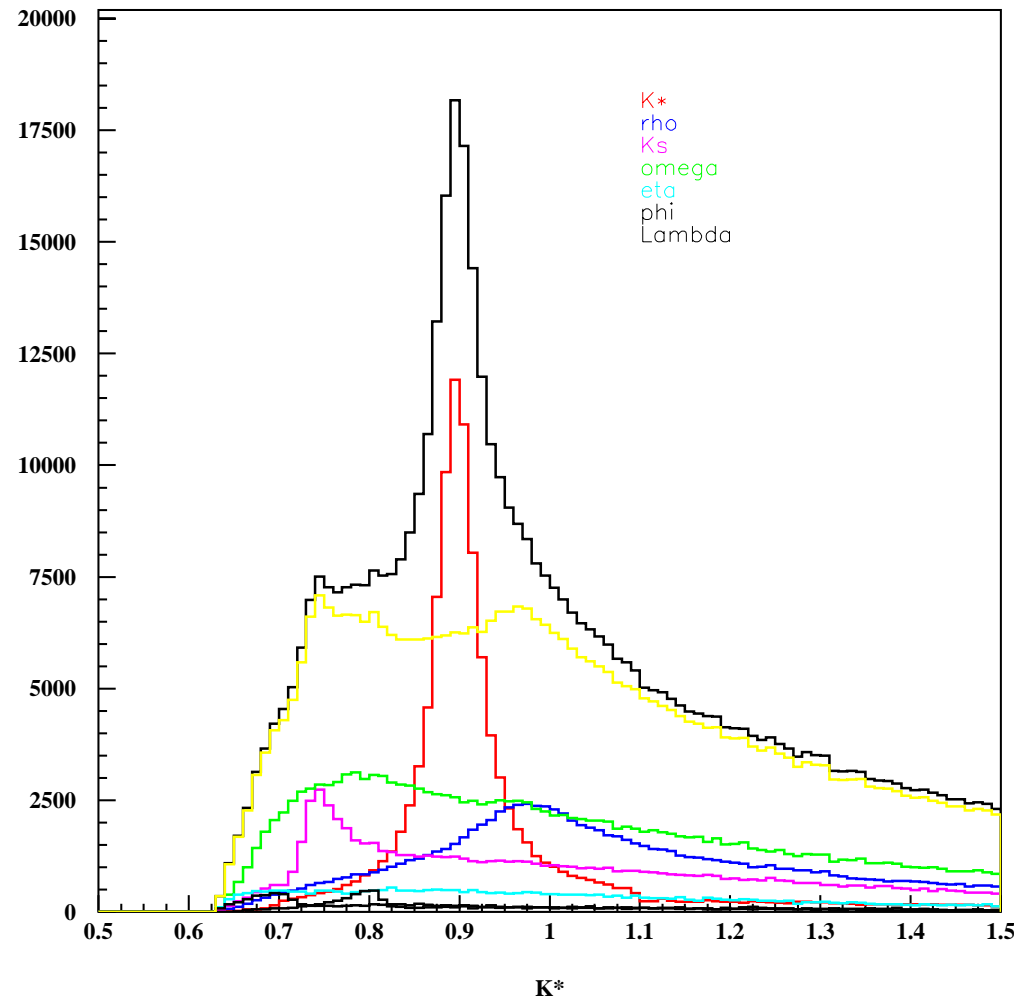


FIG. 2:  $K^{*0}$  mass plot after mixed-event background subtraction for top 14% centrally triggered events. A combination of linear background and simple Breit-Wigner resonant function is used to fit the distribution on top panel. On the bottom panel is the linear background subtracted distribution with Breit-Wigner function fit. The mass and width are found to be consistent with the standard values.

includes “partial PID” based on  $dE/dx$ ....

## Physics requiring Large-Area TOF in STAR, example 2

$\bar{d}$  and  ${}^3\bar{He}$  production in  $\sqrt{s_{NN}} = 130$  GeV Au+Au collisions

(analysis by Dave Hardtke)

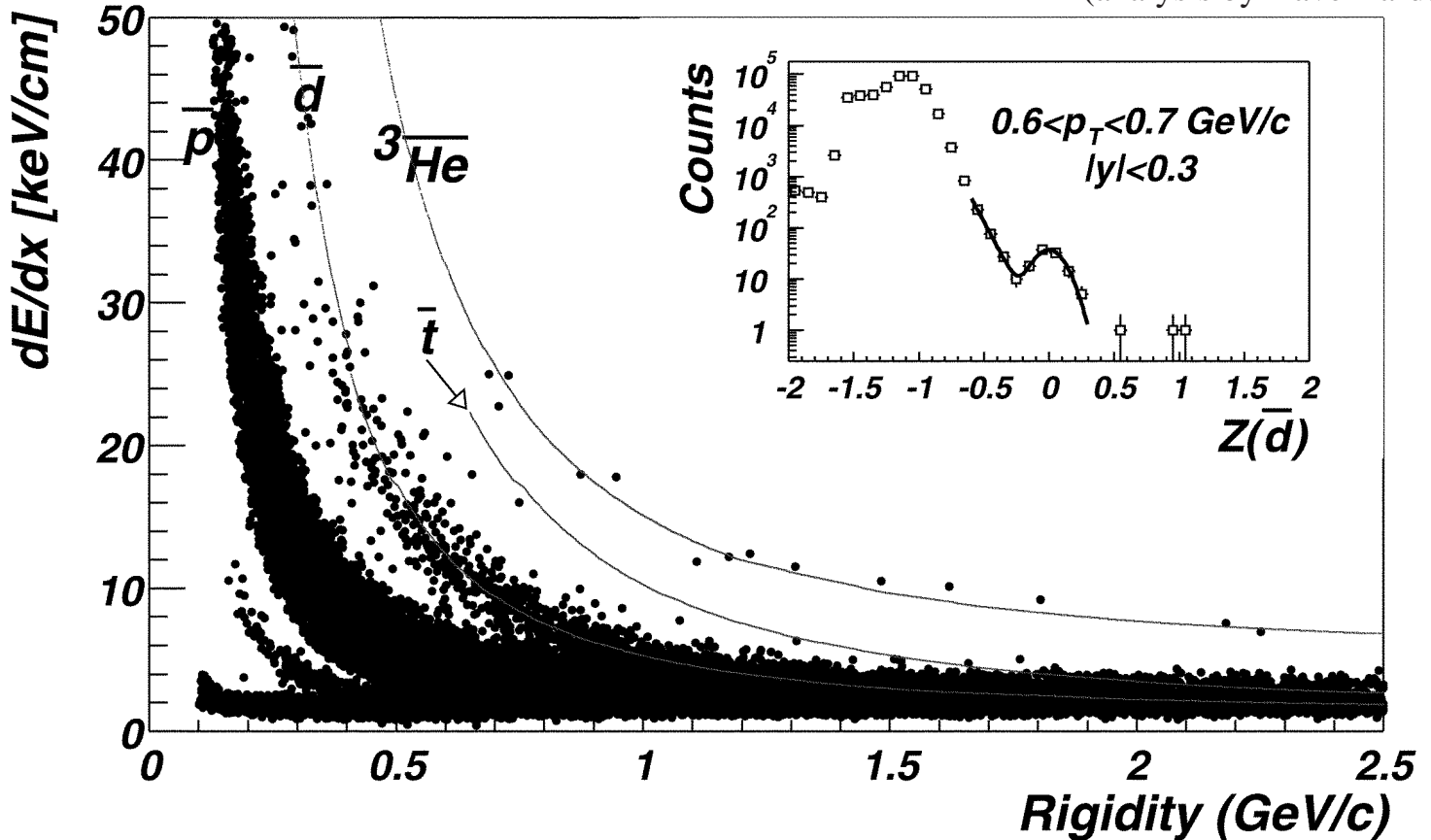


FIG. 1: Ionization ( $dE/dx$ ) versus rigidity (|momentum/nuclear charge units|) for negative tracks. The  $\pi^-$  and  $K^-$  bands have been suppressed. Also plotted are the Bethe-Bloch expectations for  $\bar{d}$ ,  $\bar{t}$  and  ${}^3\bar{He}$ . Inserted is a projection of the  $Z$  variable (see text) for one transverse momentum bin ( $0.6 < p_T < 0.7$  GeV/c).

based on  $dE/dx$  alone, STAR

- can easily identify  ${}^3\text{-He}$  (anti)fragments

- can identify only the lower momenta (anti)deuterons

- can identify only a tiny *tail* of the (anti)triton spectrum

all three bands (and the important “reference”  $p$  and  $pbar$  bands) are well-separated out to much higher momenta with TOF information





# Resistive Plate Chambers

(in avalanche mode)

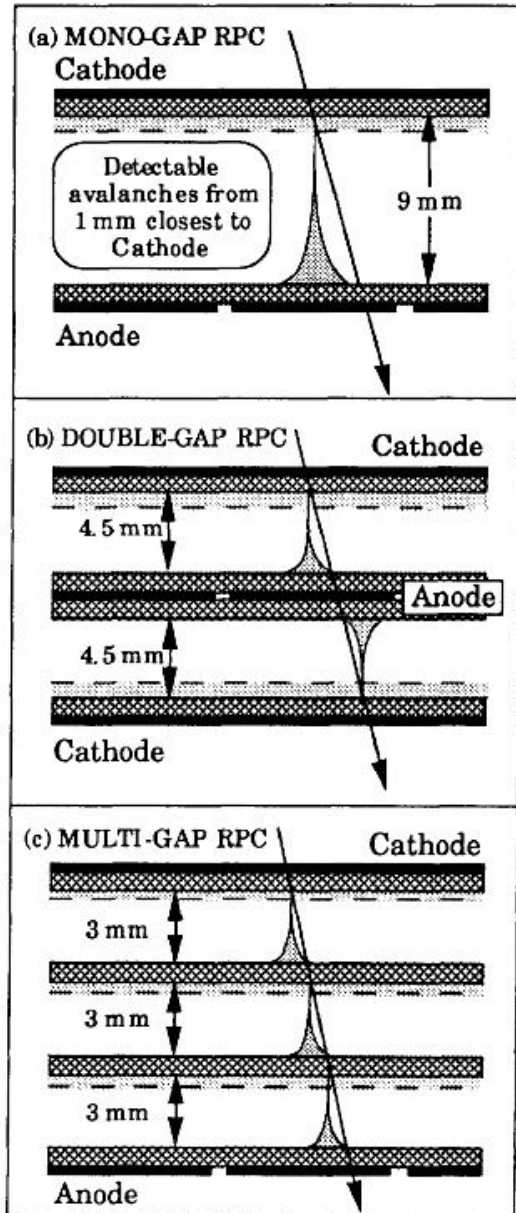


Figure 2. Schematic representation of a monogap RPC (top); a double gap RPC (middle) and a multigap RPC (bottom).

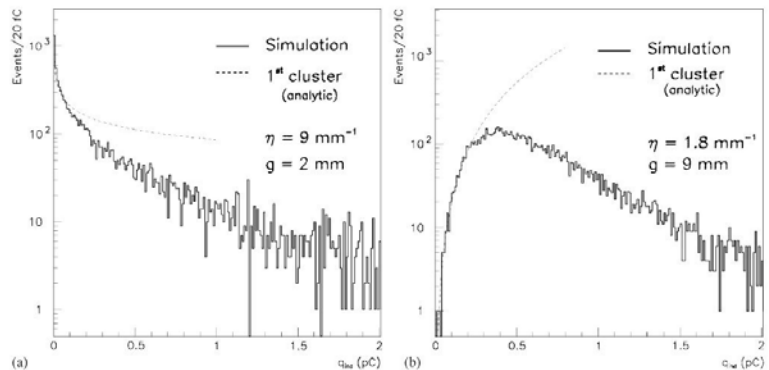


Fig. 1. Comparison between the simulated charge spectra for single (a) narrow ( $g = 2$  mm) and (b) wide ( $g = 9$  mm) gap RPCs.

narrow single gaps don't work well in avalanche mode

wider single gaps?

enhanced streamer-free range of operating voltage  
but time resolution suffers...

- primary ionization is a stochastic process!  
→ timing jitter from location of ionization in RPC
- avalanches from single primary clusters tend to merge  
→ fluctuations in avalanche development dominate

many narrow gaps!

- characteristic distance for primary ionization decreased  
→ decreased timing jitter from primary ionization step
- N-independent avalanches, hence an averaging  
→ decreased timing jitter from avalanche fluctuations

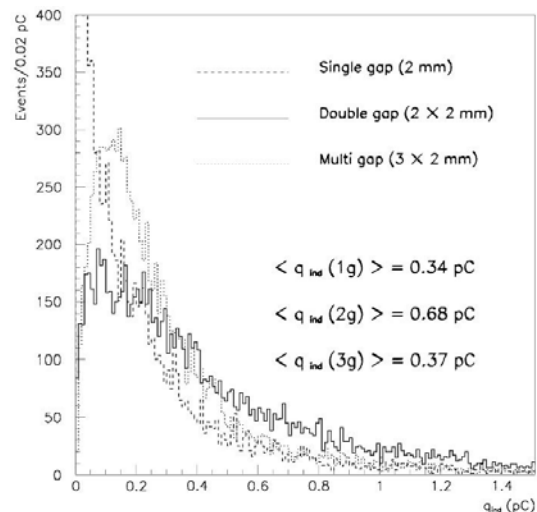


Fig. 3. Monte Carlo charge spectra for single- double- and multi-gap RPCs.

With many small gaps,

- the resistive plates are transparent to the fast signals generated by the avalanches in each gas gap. The induced signal is the *sum over all gaps*.

- the internal resistive plates are electrically floating, so the stable state is *equal gain in all gaps*.

(ALICE prototype, figures from M. Spiegel, NIM A 453, 308 (2000)).

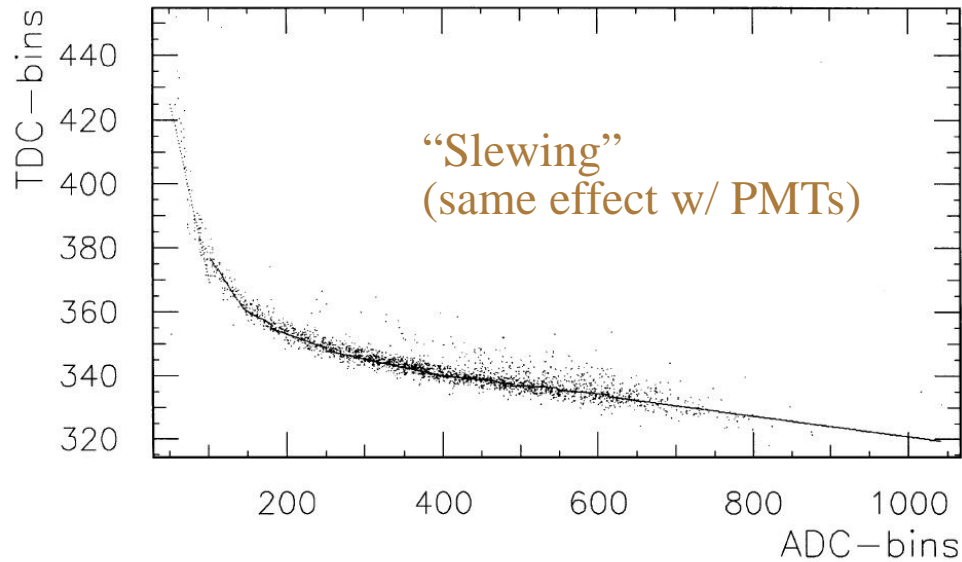


Fig. 4. Correlation of raw time difference between MRPC detector and reference counter and MRPC signal amplitude.

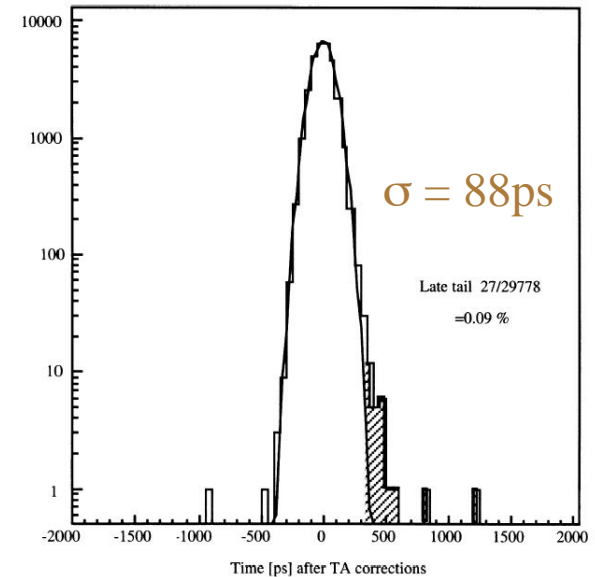


Fig. 5. Time distribution measured with an MRPC with five gaps of 220  $\mu\text{m}$  operated at 12 kV (109 kV/cm). The width of the Gaussian fit is  $\sigma = 88$  ps.

For Additional Details... (articles in blue are “must reads”!)

- “The development of the multigap resistive plate chamber”, M.C.S. Williams, Nucl. Phys. B (Proc. Suppl.) **61B**, 250 (1998).
- “Microstreamers and the micro-gap Resistive Plate Chamber”, E. Cerron-Zeballos *et al.* (inc. Lamas-Valverde, Platner, & Roberts), NIM A **411**, 51 (1998)
- “Effect of adding SF6 to the gas mixture in a multigap resistive plate chamber”, E. Cerron-Zeballos *et al.* (inc. Lamas-Valverde, Platner, & Roberts), NIM A **419**, 475 (1998)
- “A very large multigap resistive plate chamber”, E. Cerron-Zeballos *et al.* (inc. Lamas-Valverde, Platner, & Roberts), NIM A **434**, 362 (1999)
- “Recent Progress on RPCs for the ALICE TOF system”, M. Spiegel for the ALICE Collaboration, NIM A **453**, 308 (2000)
- “A new high-resolution TOF technology”, P. Fonte *et al.*, NIM A **443**, 201 (2000)
- “A Multigap Resistive Plate Chamber prototype for the Time-Of-Flight of STAR at RHIC”, J. Lamas-Valverde *et al.*, Vienna Conference on Instrumentation, 2001 (submitted)
- “Optical observation of discharge in resistive plate chamber”, I. Kitayama *et al.*, NIM A **424**, 474 (1999)
- “Test beam results on resistive plate chamber prototype at gamma irradiation facility in CERN”, C.H. Chung *et al.*, NIM A **432**, 14 (1999)
- “Study of avalanche mode operation of resistive plate chambers with different gas gap structures”, V.V. Ammosov *et al.*, NIM A **441**, 348 (2000)

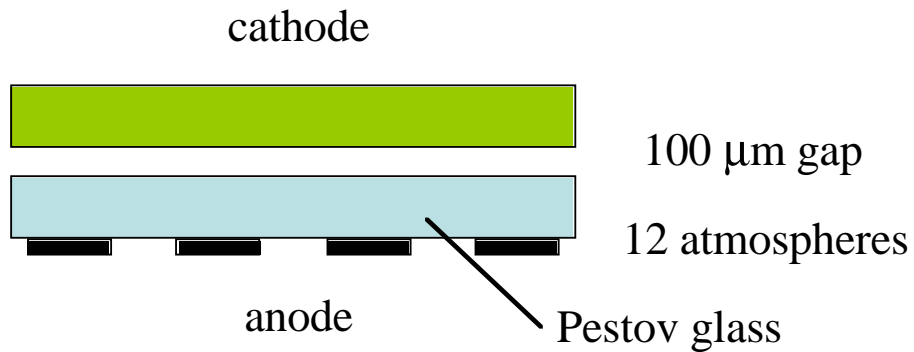
...and on the theory side...

- “Resistive Plate Chambers in avalanche mode: a comparison between model predictions and experimental results”, M. Abbrescia *et al.*, NIM A **409**, 1 (1998)
- “A simple theory of signals induced by a point charge moving in a resistive plate chamber”, Th. Heubrandtner *et al.*, NIM A **419**, 721 (1998)
- “The simulation of resistive plate chambers in avalanche mode: charge spectra and efficiency”, M. Abbrescia *et al.*, NIM A **431**, 413 (1999)

## Two gaseous detectors considered for ALICE

### Pestov counters

Glass electrode and metal electrode



Excellent time resolution  $\sim 50$  ps

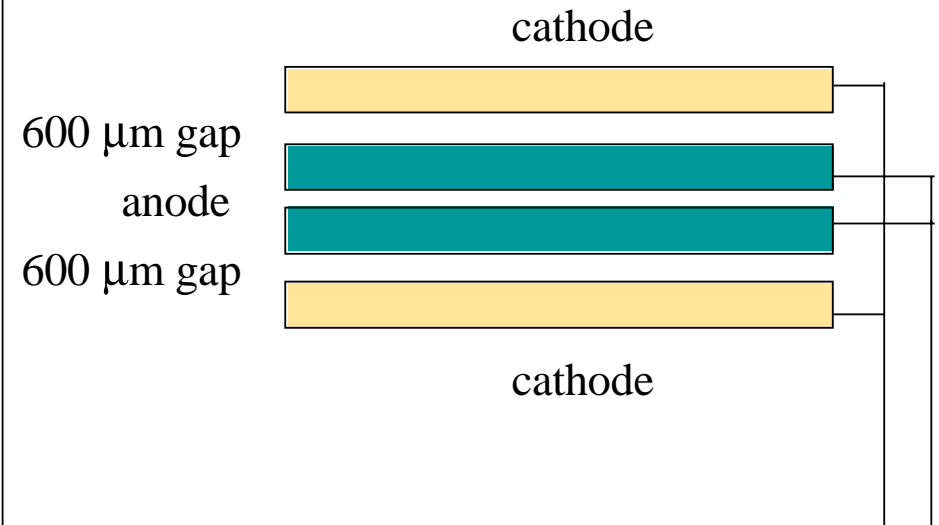
But long tail of late events

Mechanical constraints (due to high pressure)

Non-commercial glass

### Double gap PPC

Both electrodes metallic



Marginal time resolution  $\sim 250$  ps

Small signal (to keep sparks at low probability)

Difficult to build

○ P. Fonte et al. / Nuclear Instruments and Methods in Physics Research A 443 (2000) 201 }204

✕ V.V. Parchomchuck, Yu.N. Pestov, N.V. Petrovykh, Nucl. Instr. and Meth. 93 (1971) 269.

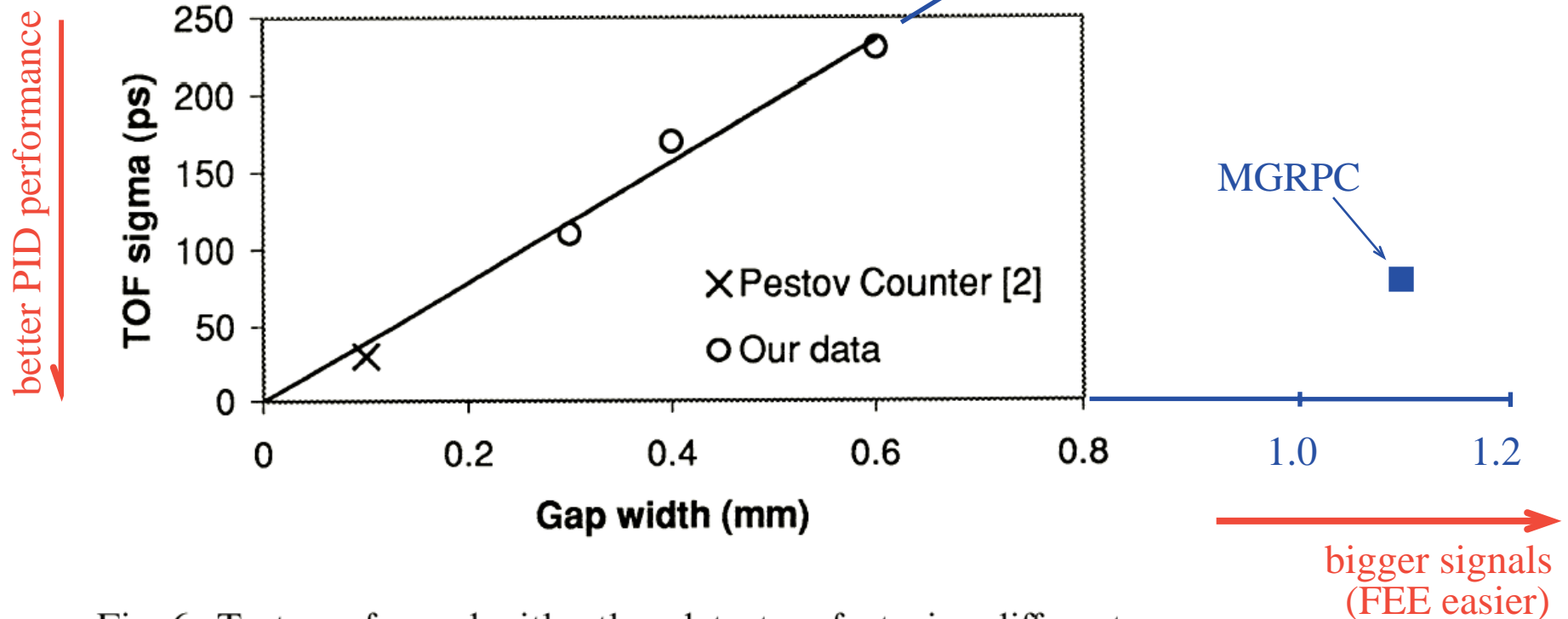
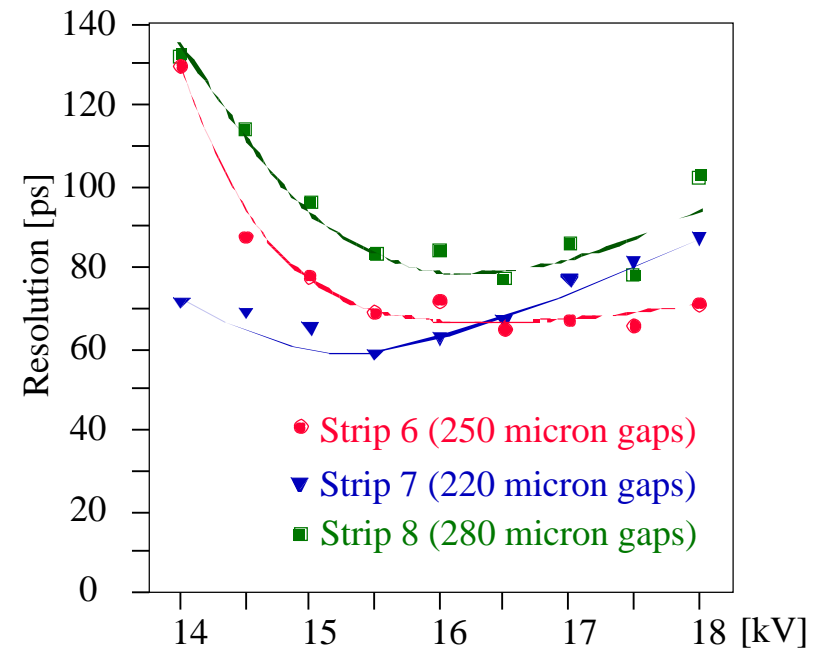
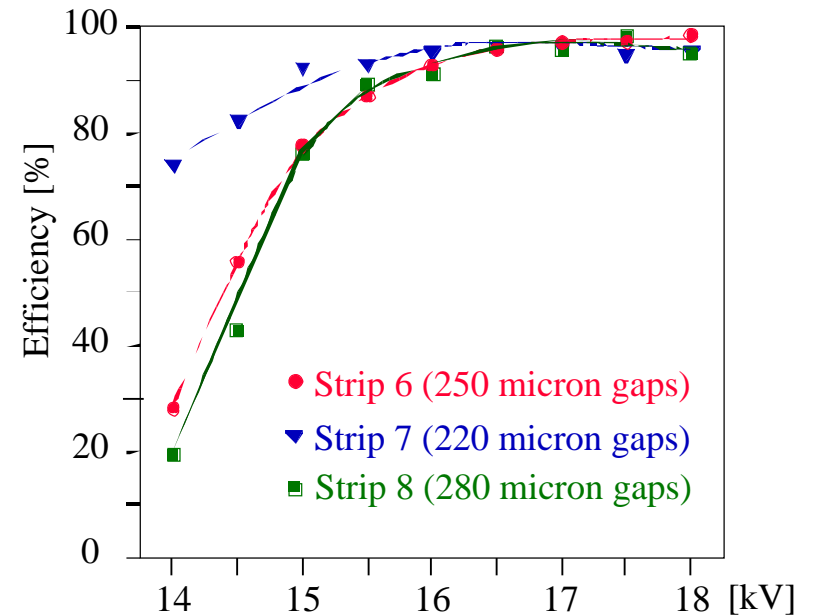


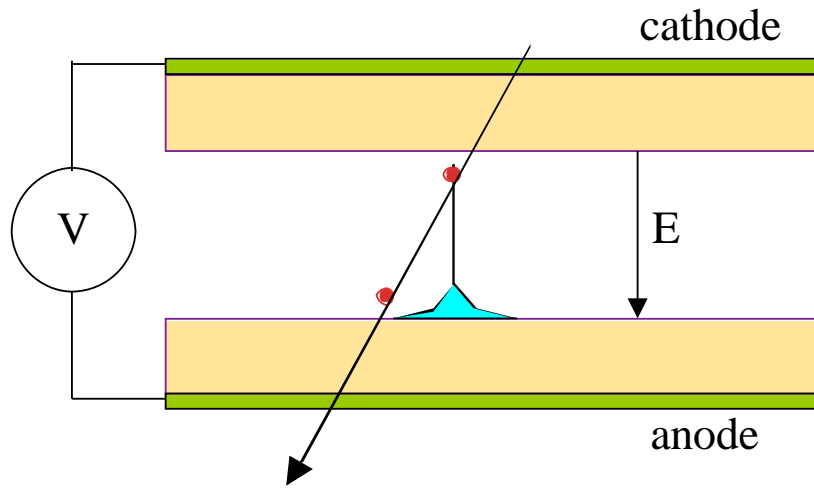
Fig. 6. Tests performed with other detectors featuring different widths of the gas gap suggest that the main contribution to the time jitter is associated to the amplification process in the gas. The timing resolution seems to depend almost linearly on the gap width, with a slope of approximately 40 ps/0.1 mm.

MGRPCs are (relatively) easy to build!

Big change in gap size  $\rightarrow$  small change in operating voltage. Large 'plateau' region where efficiency high, time resolution excellent and gap can vary by  $\pm 30 \mu\text{m}$

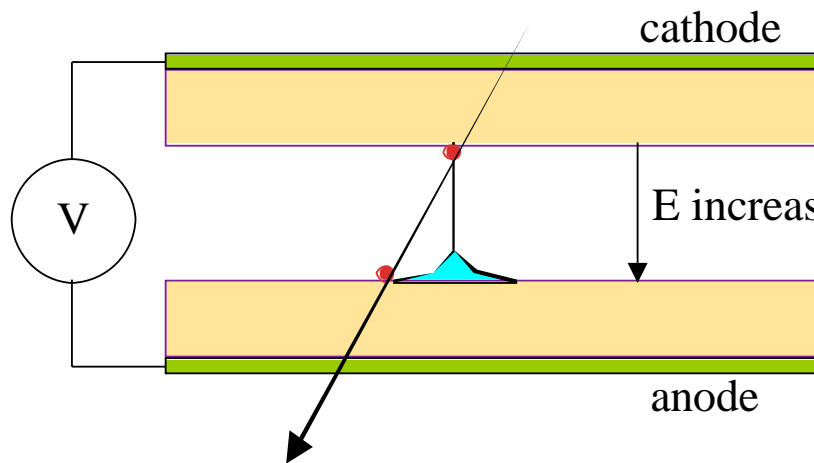
Thus device with this excellent time resolution can be built with very 'relaxed' mechanical tolerances





Charged particle passes through gas gap and creates clusters of electrons and positive ions electrons avalanche towards anode → fast signal on external electrodes - etc

Now consider smaller gap



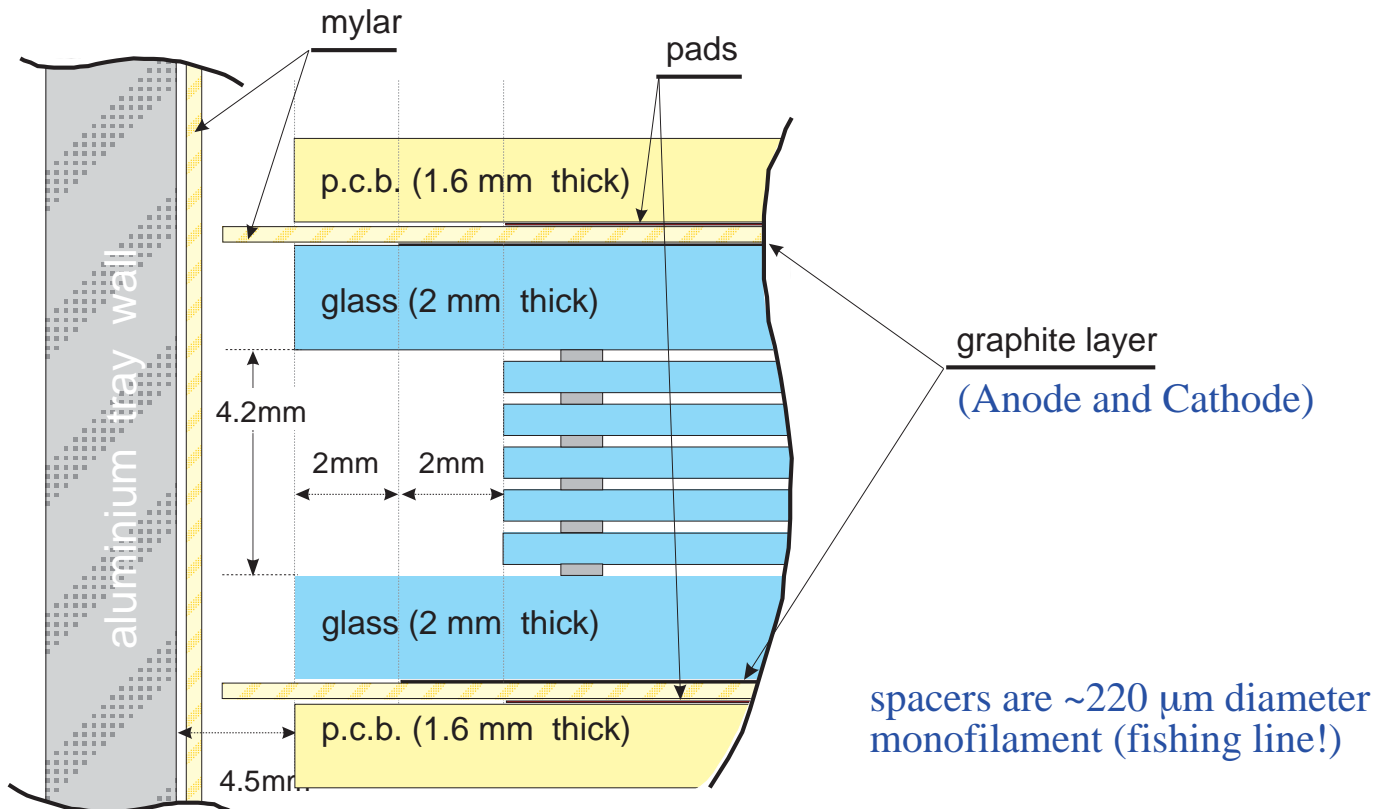
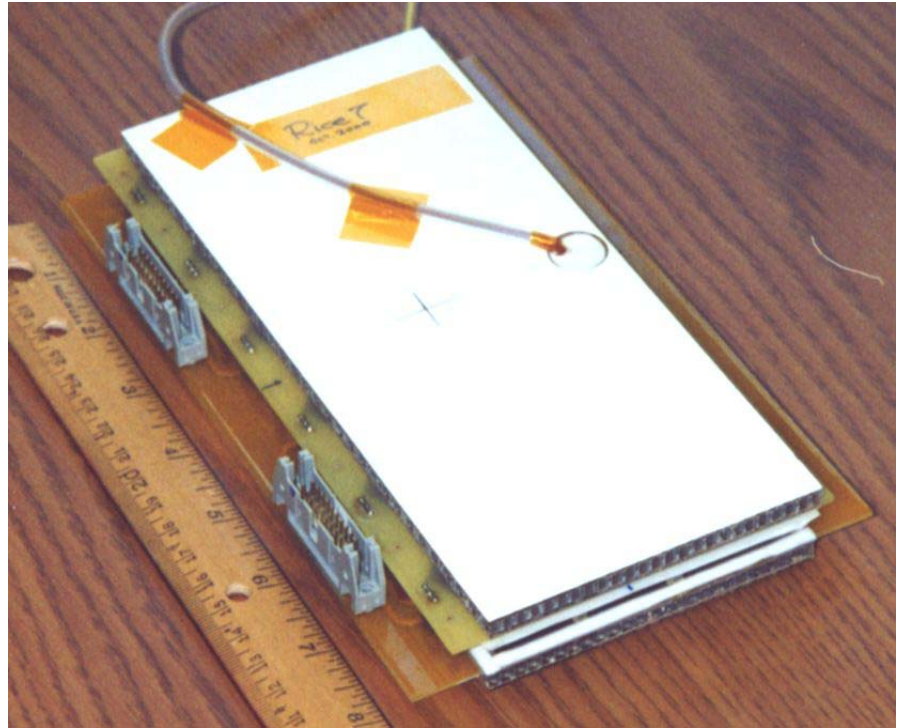
Thus **Townsend coefficient higher** - bigger avalanche (i.e. **higher gain**)  
 however gap smaller therefore **less distance for avalanche to grow** (i.e. **lower gain**)

Apparently we are working in region where both effects cancel (by 'magic' it is rather an exact cancellation)

## Rice Version 7 MGRPC

...costs **~30\$/ch...**

...compared to **~1700\$/ch**  
for equiv. Conventional TOF...



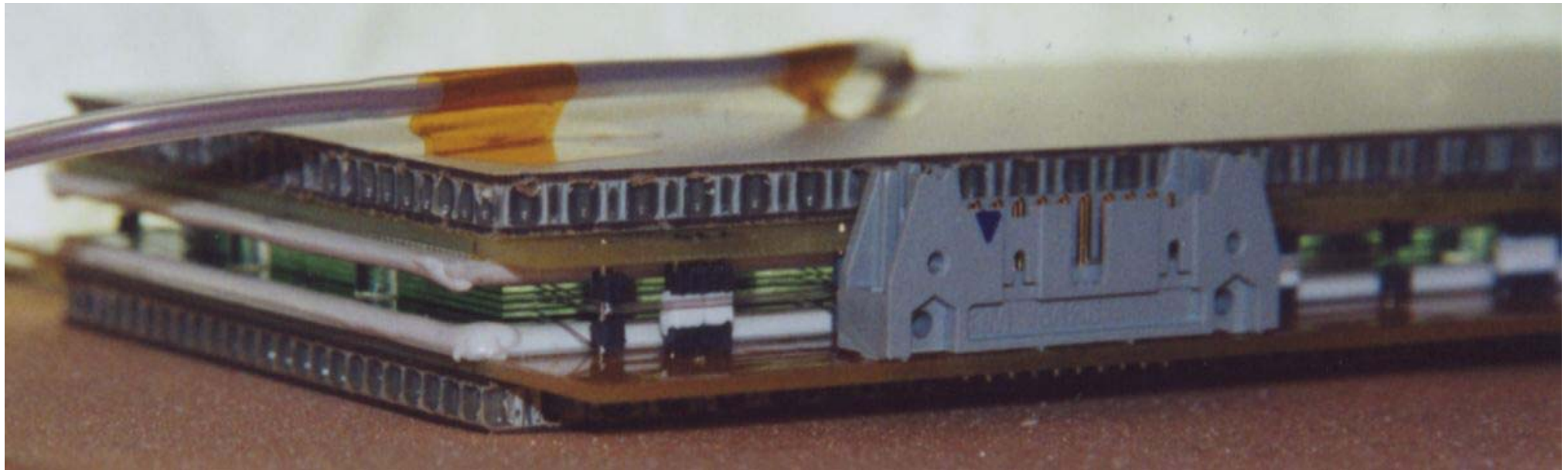
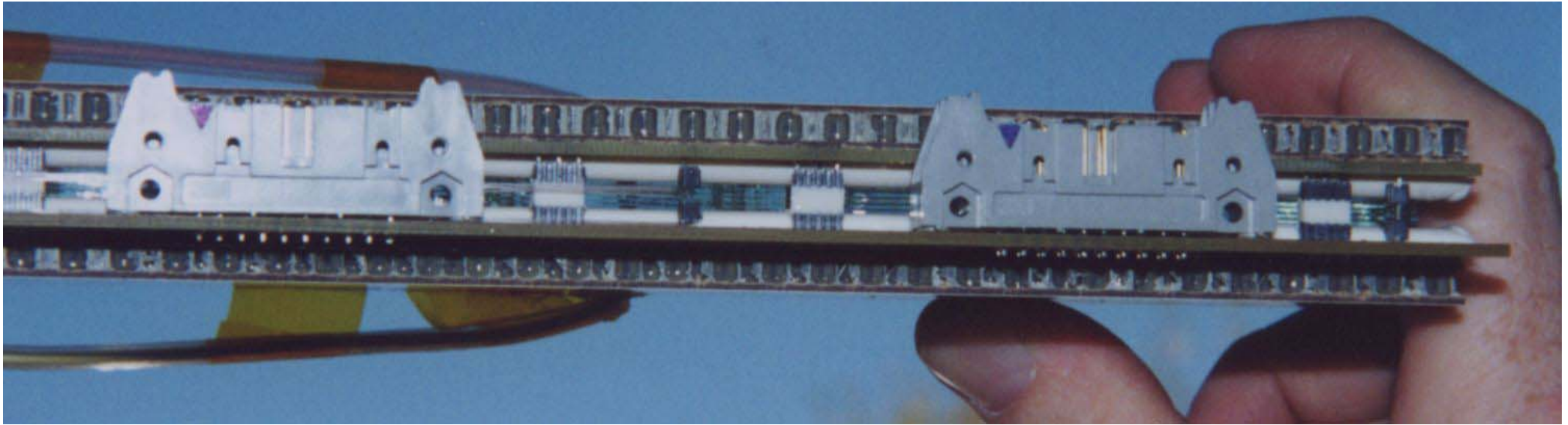
Tray inner dimensions: 213mmx2413mmx85mm

outer glass: 84mm x 204mm (10mm extra width for HV / gnd. connect. included)

graphite layer: 70mm x 200mm

inner glass: 65mm x 195mm

pad size: 31mm x 30mm, distance between pads = 3mm



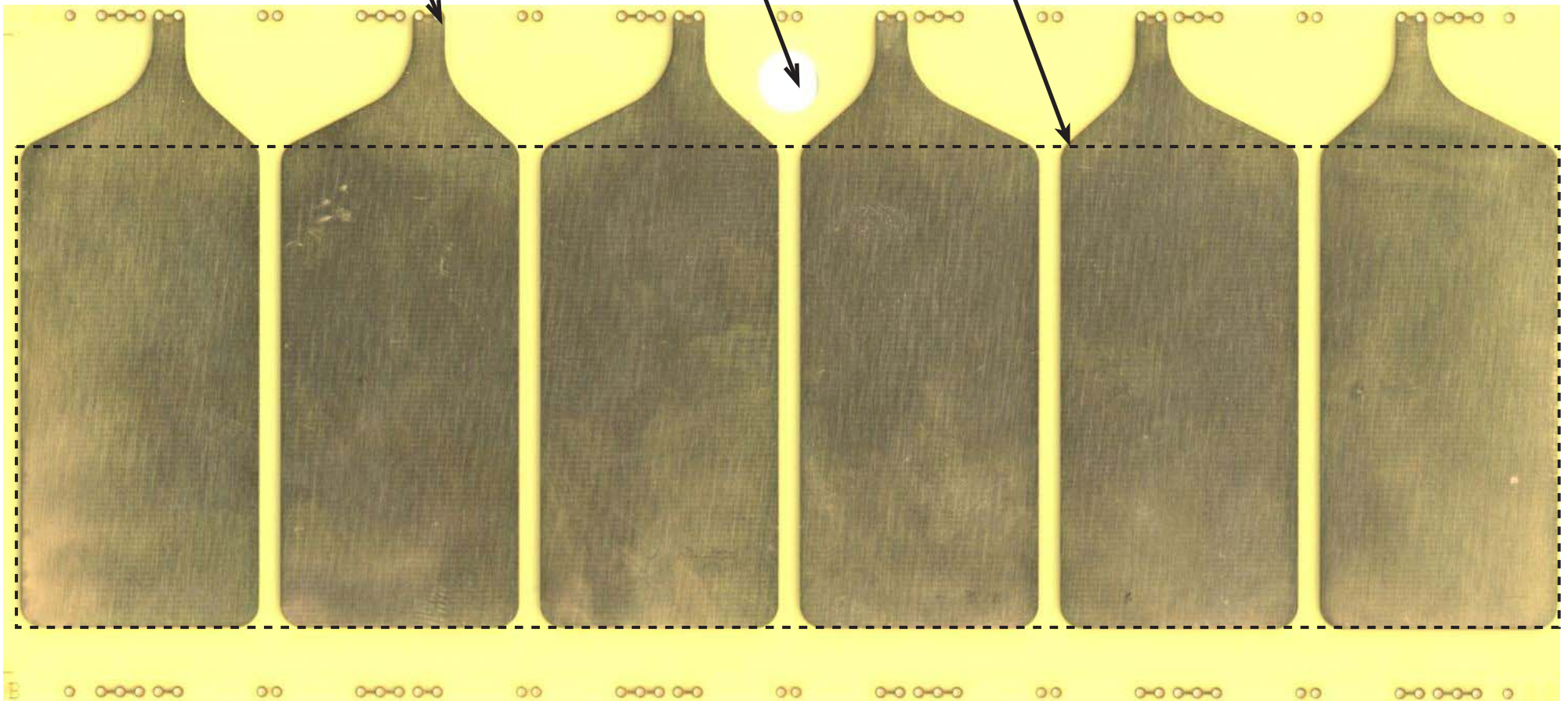


## Readout Pads

region of pads above thin glass indicated by dashed box

hole to pass HV through to graphite layer

pins for signal output



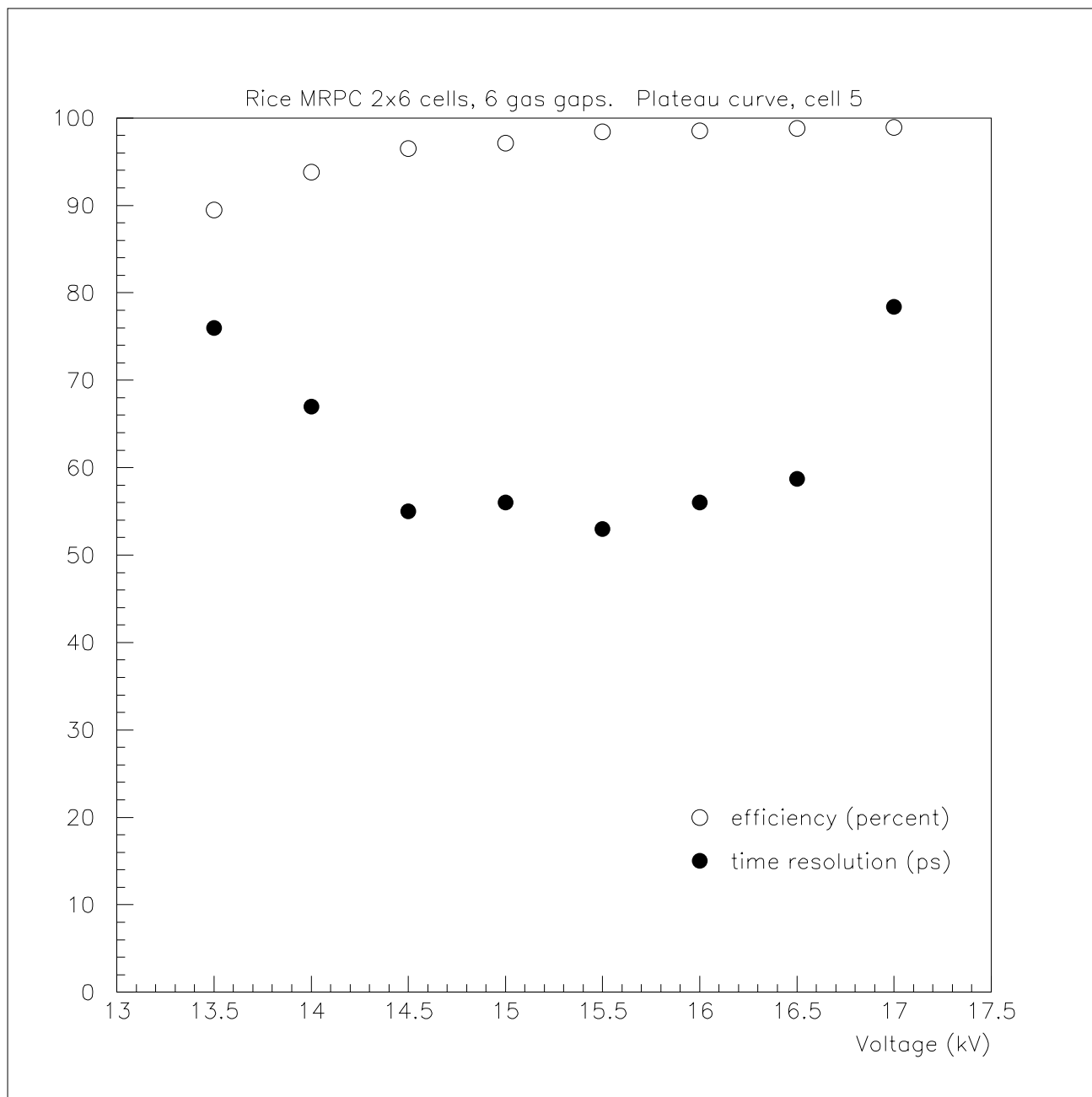


Fig. 2. The efficiency and corrected time resolution ( $\sigma$ ) versus voltage for the prototype with 2x6 pads and pad area of 3x3 cm<sup>2</sup>.

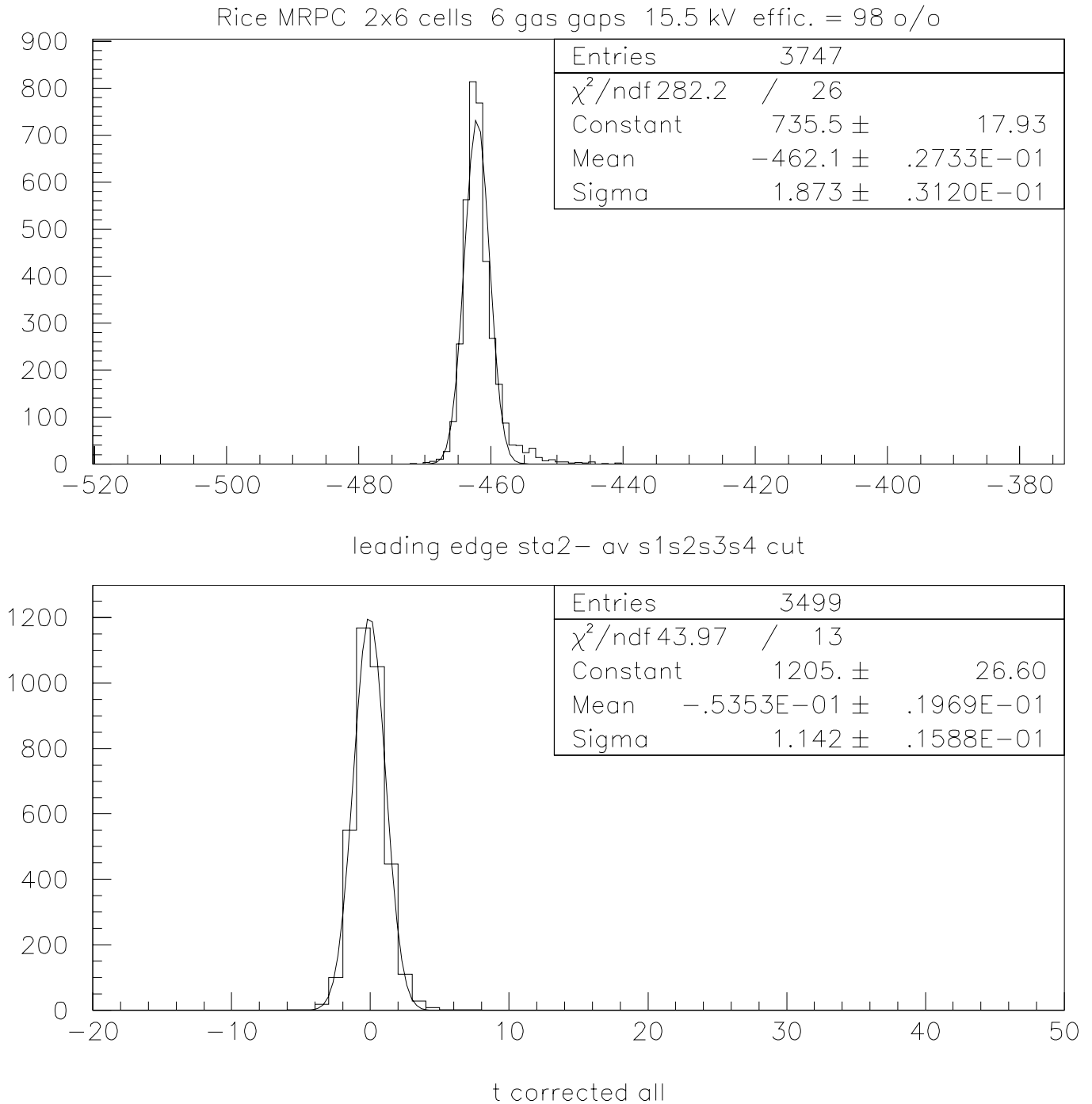
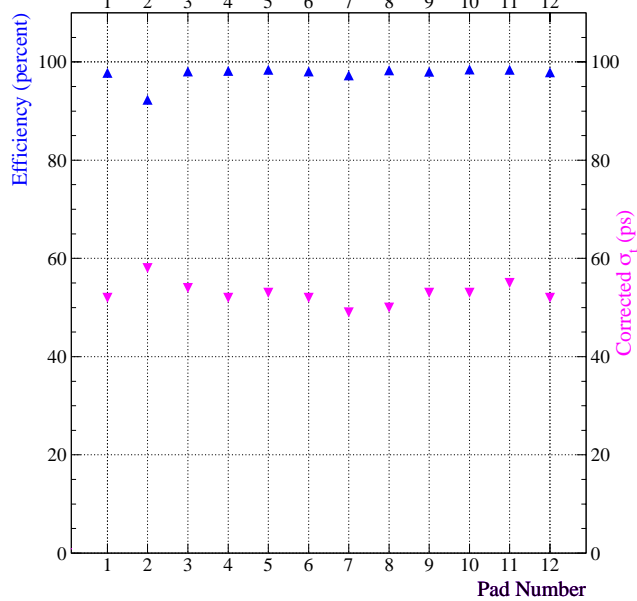


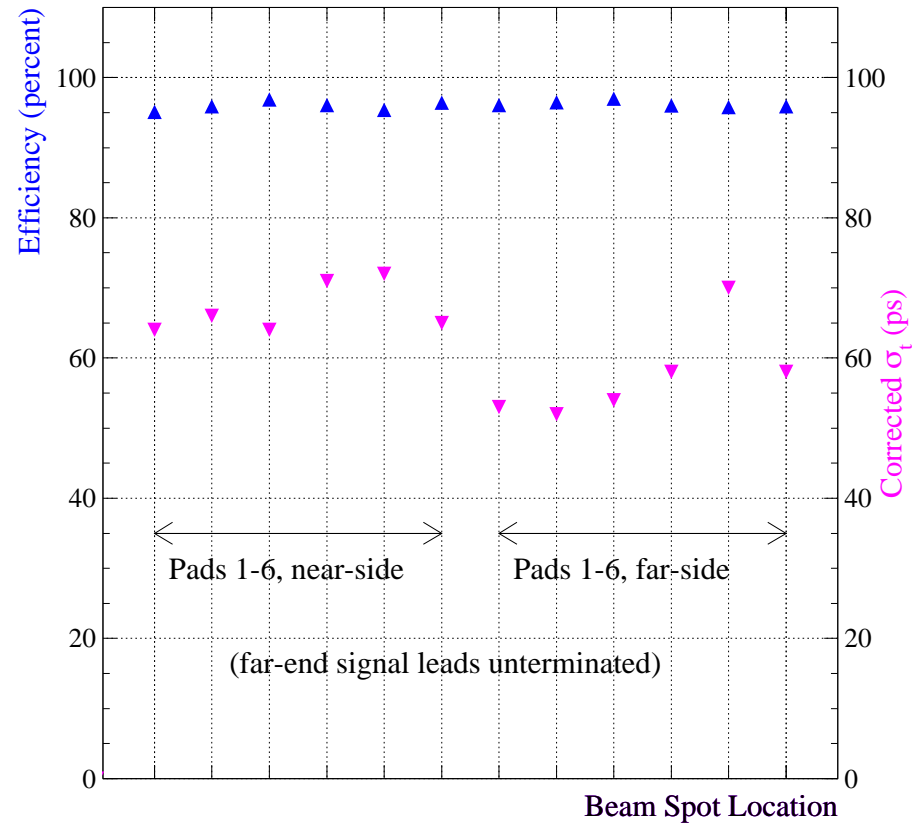
Fig. 1. The histogram on the top shows the raw time, and the histogram on the bottom shows the Time-to-Amplitude corrected time. The distributions are fitted with Gaussians, and the variance,  $\sigma$ , is reported in the boxes in each frame in units of TDC bins, which are 50 ps wide. Comparing the two frames, one observes the overall timing resolution improves from 94 ps to 57 ps following the offline correction for slewing [3] and, by subtracting the mean jitter introduced by scintillators counters, we obtain 90 ps and 50 ps respectively.

# Channel-by-Channel Variations

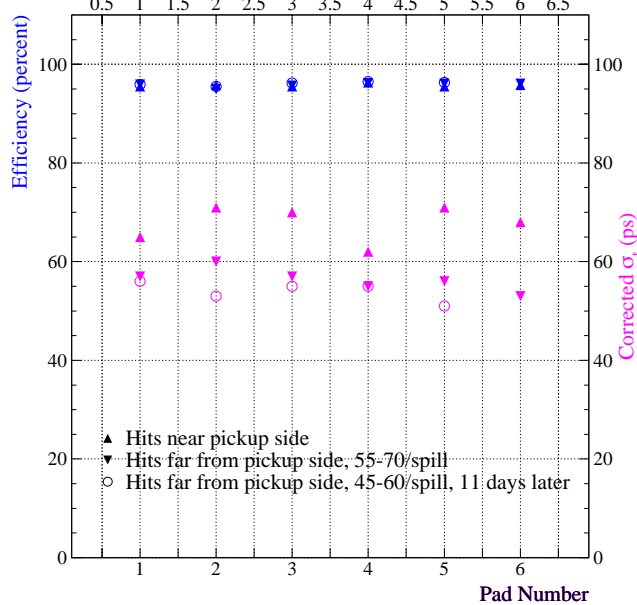
Rice version 6, 2x6, 3x3.1cm, 15.5 kV, Fall 2000 Test beam



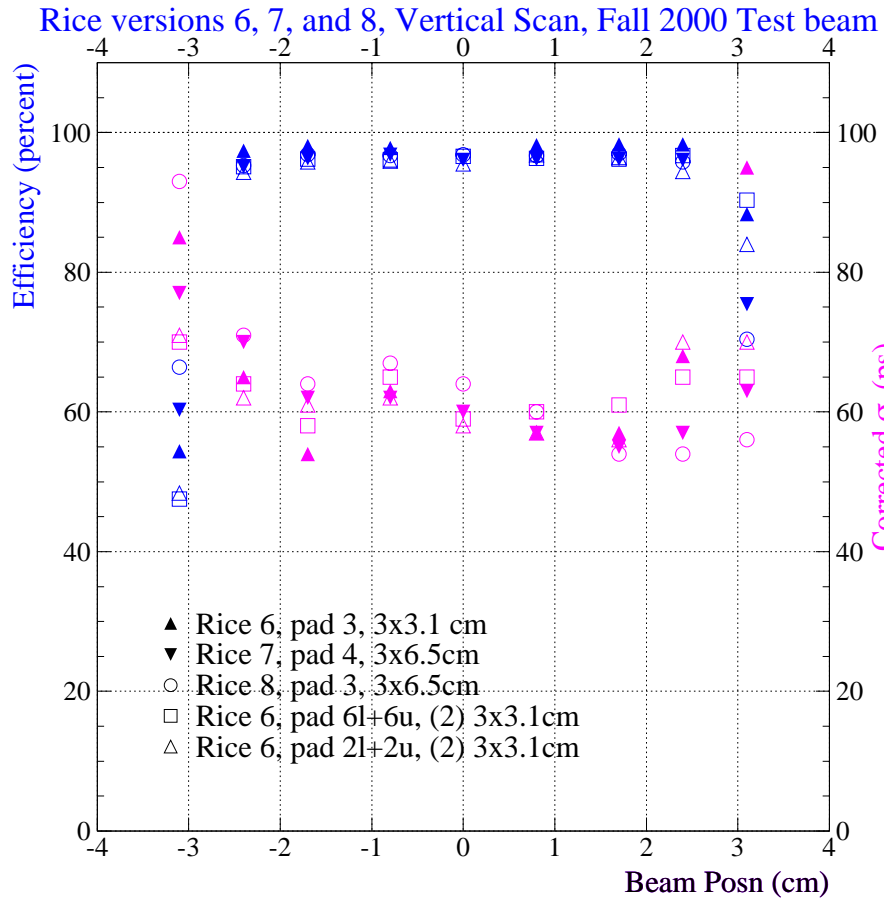
Rice version 8, 1x6, 3x6.5cm, 15.5 kV, Fall 2000 Test beam



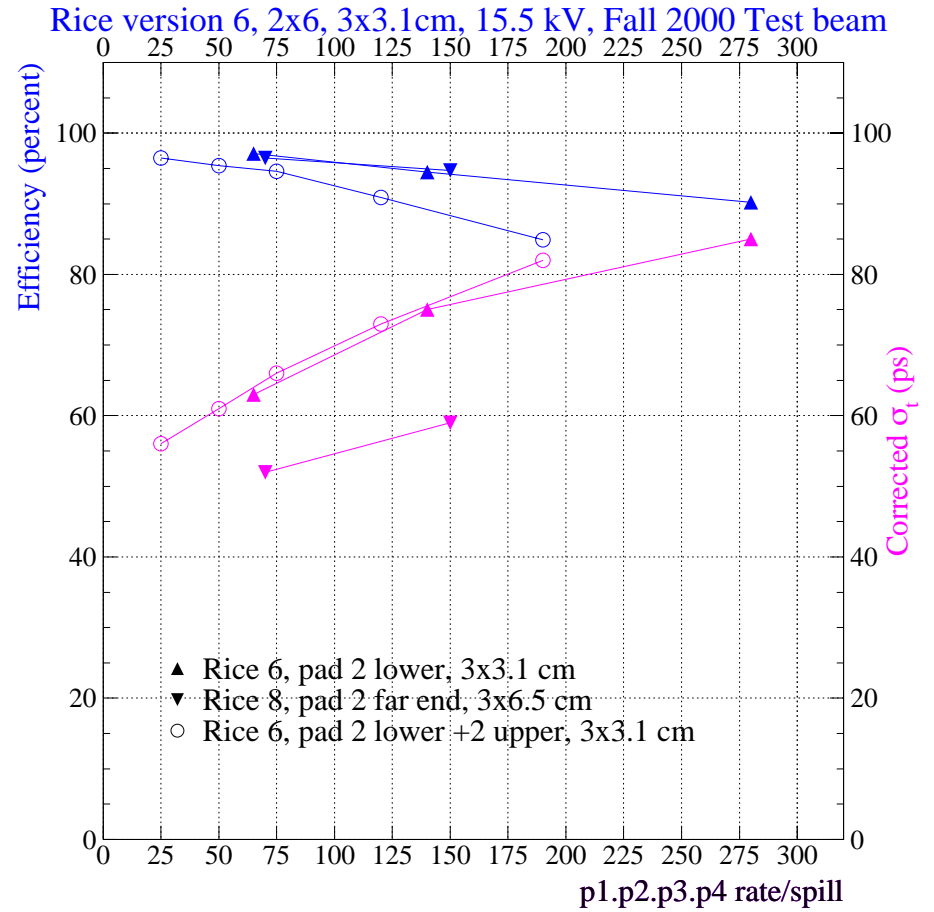
Rice version 7, 1x6, 3x6.5cm, 15.0 kV, Fall 2000 Test beam

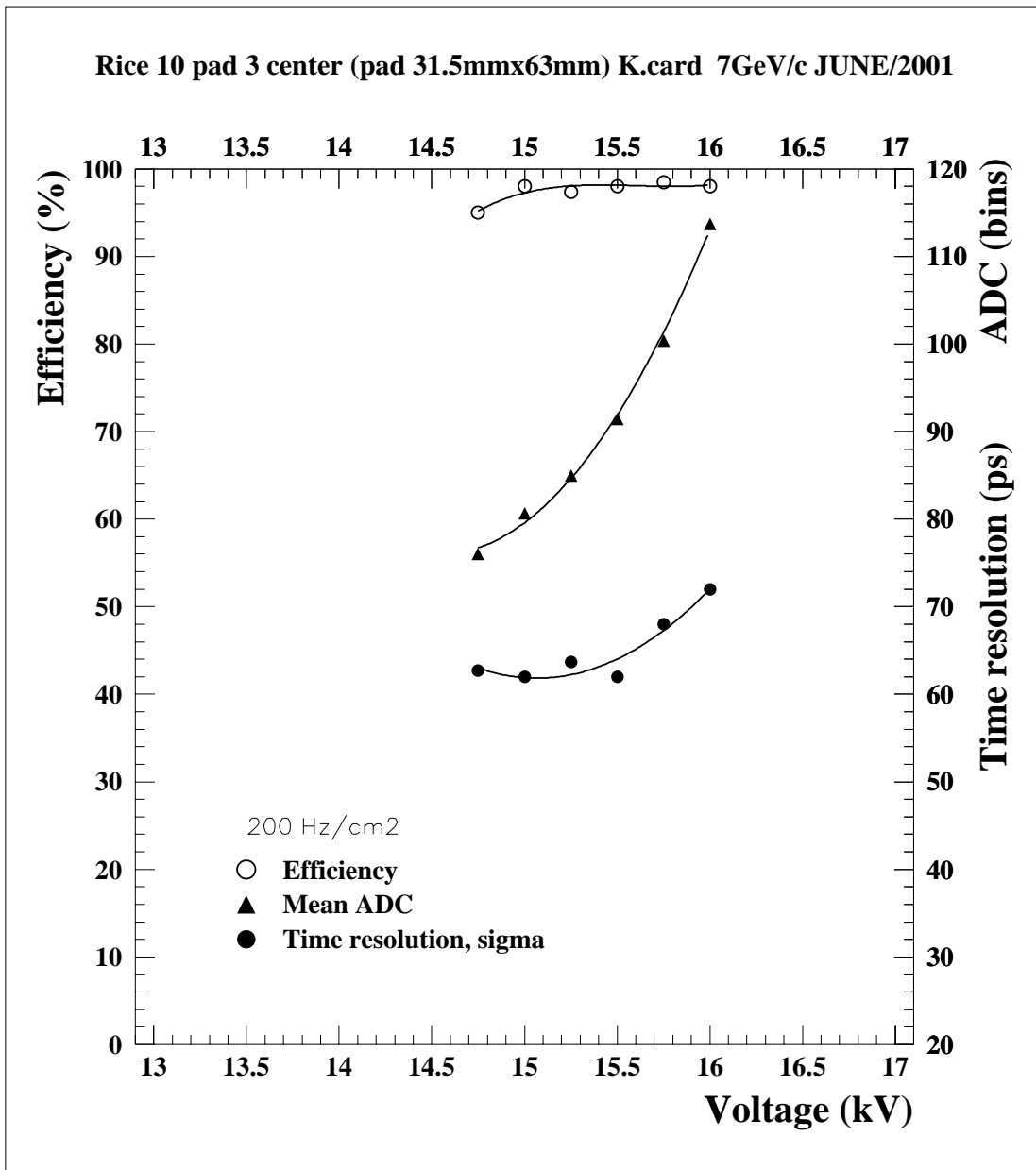


# Position Dependence w.r.t. Pads



# Rate Dependence





Rice versions 6-8  
good & uniform & reproducible performance....

Rice versions 9, 10, ...  
no more changes to MGRPC interior structure (glass, gaps, ...)  
only optimization (minimization of dead space etc....)

# PROPOSAL

To the

## STAR Collaboration

For the construction of a

# Single Tray of Multi-gap Resistive Plate Chamber Time-of-Flight for STAR

By the STAR TOF group\*

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Others TBA

# Front-End Electronics

typical MIP hit in a MGRPC equivalent to  $\sim 25$  fC (not a typo)

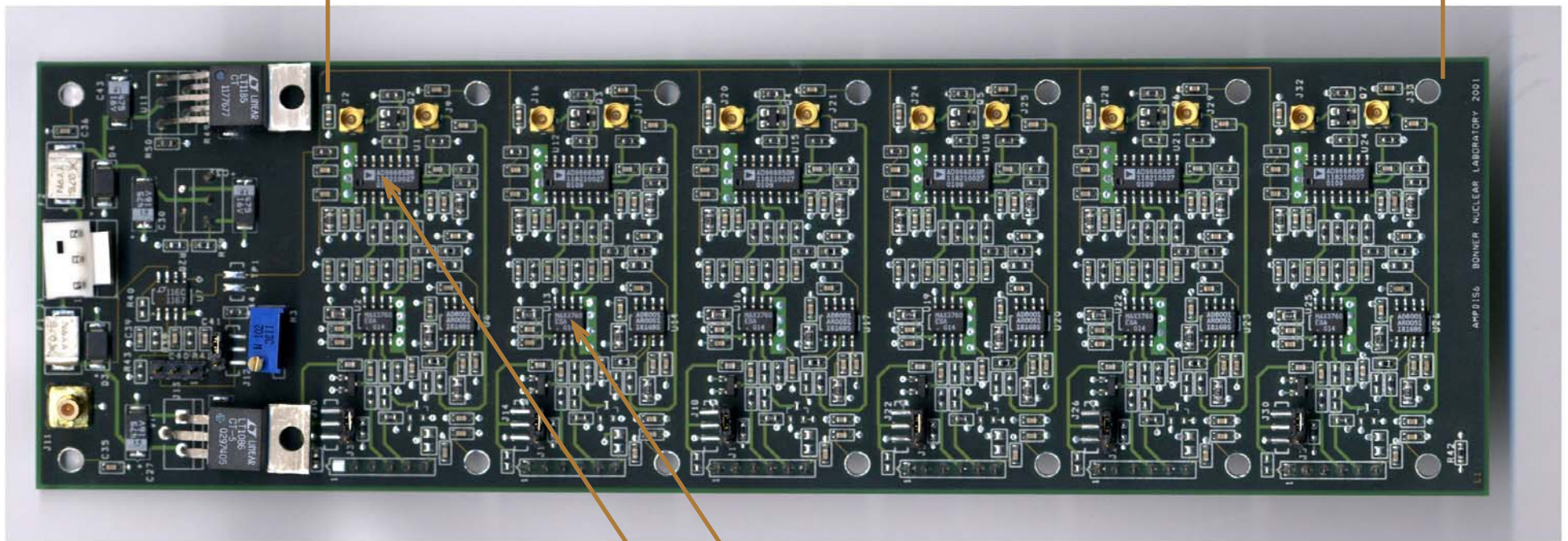
need to amplify first

rise time of [detector+FEE] practically limited only by bandwidth of this preamp

major breakthrough for ALICE and STAR FEE development came w/ adoption of MAXIM 3760

then discriminate using standard components

6 channels, one per MGRPC pad

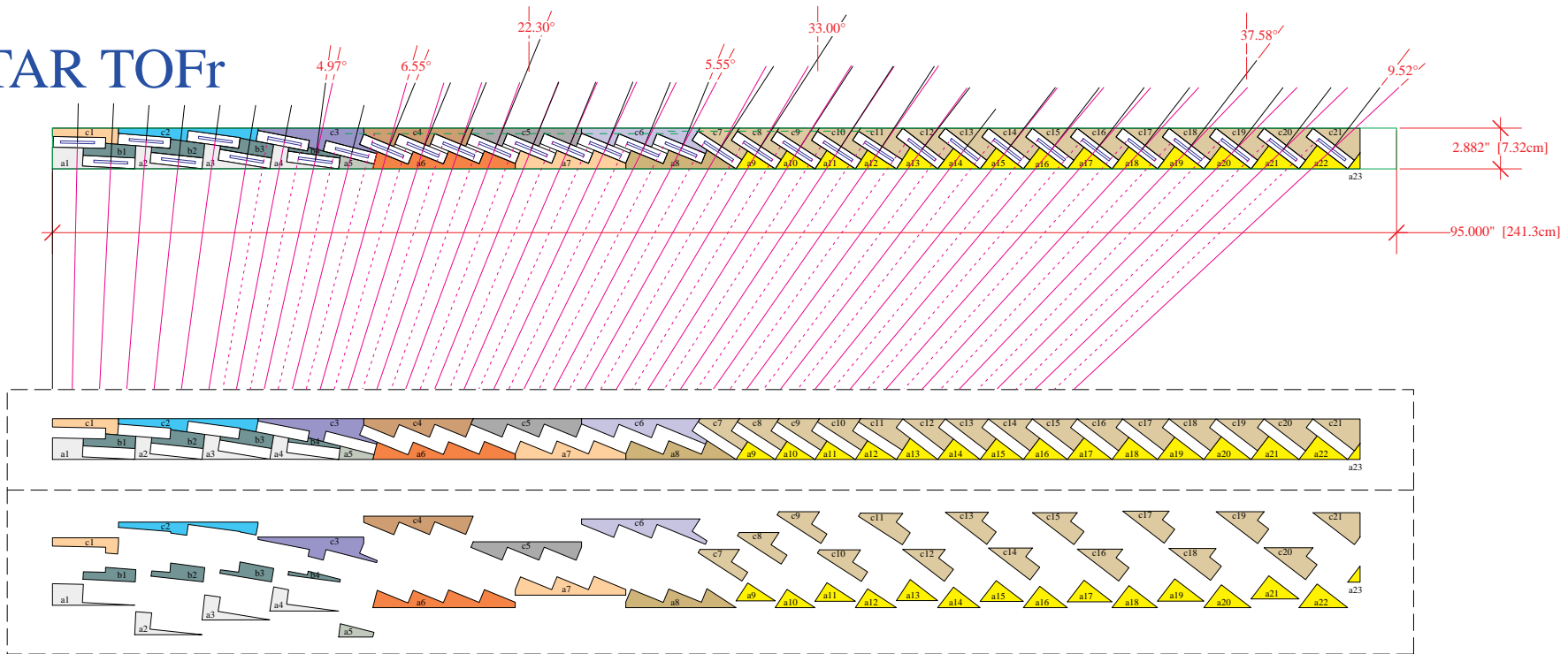


Maxim 3760 Preamplifier

Analog Devices 96687 Comparator (TOFp, pVPD, TOFr)



# STAR TOFr

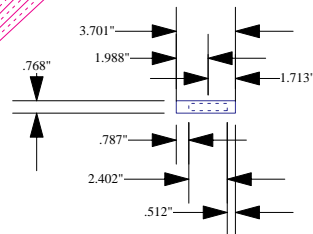


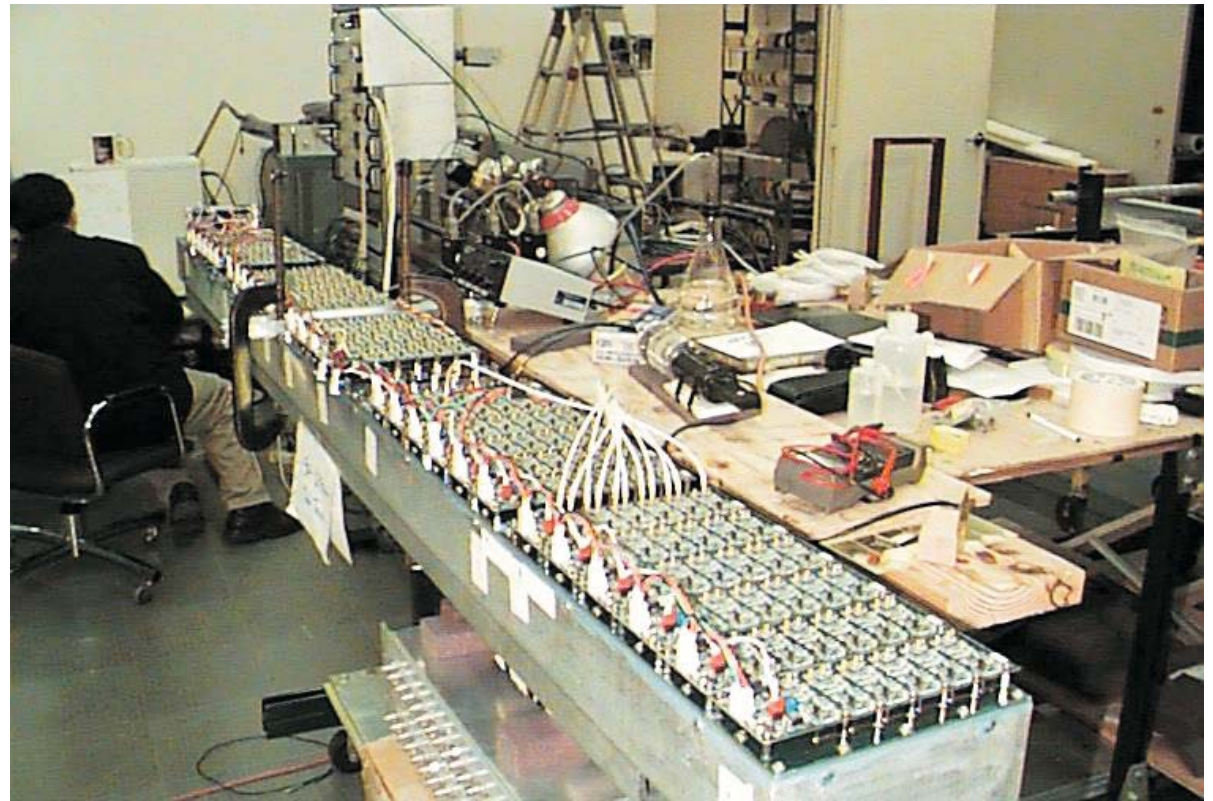
## A Production Prototype for Large-Area TOF for STAR

- 28 MGRPC modules fabricated at USTC and CERN
- 33 FEE boards fabricated at Rice
- Tray construction at Rice
- Construction completed February 16, 2002
- under test at Rice, soon to be at BNL for test beams...

Installation STAR for next RHIC run (Fall 2002)

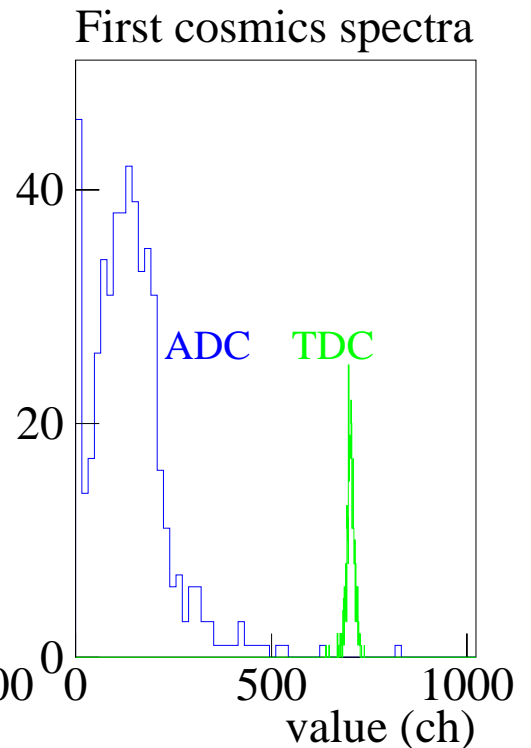
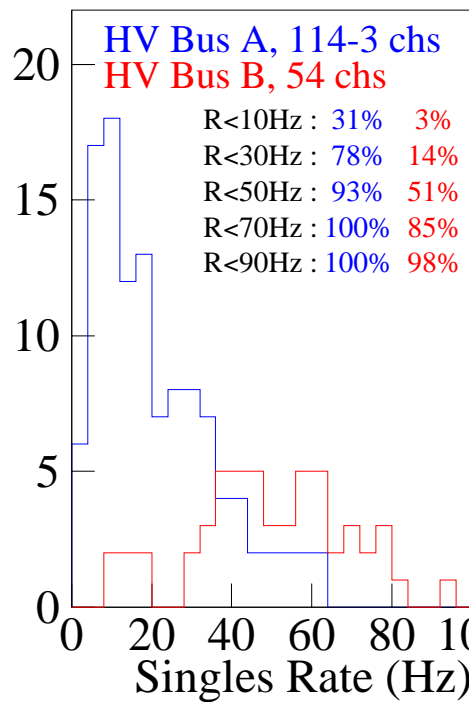
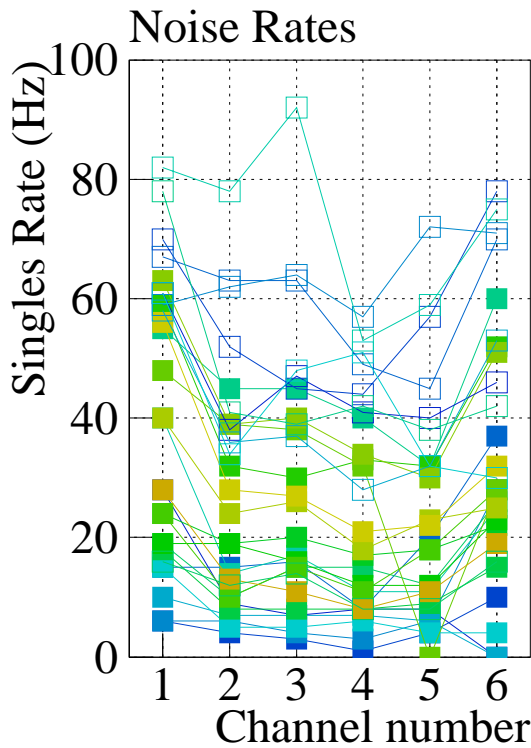
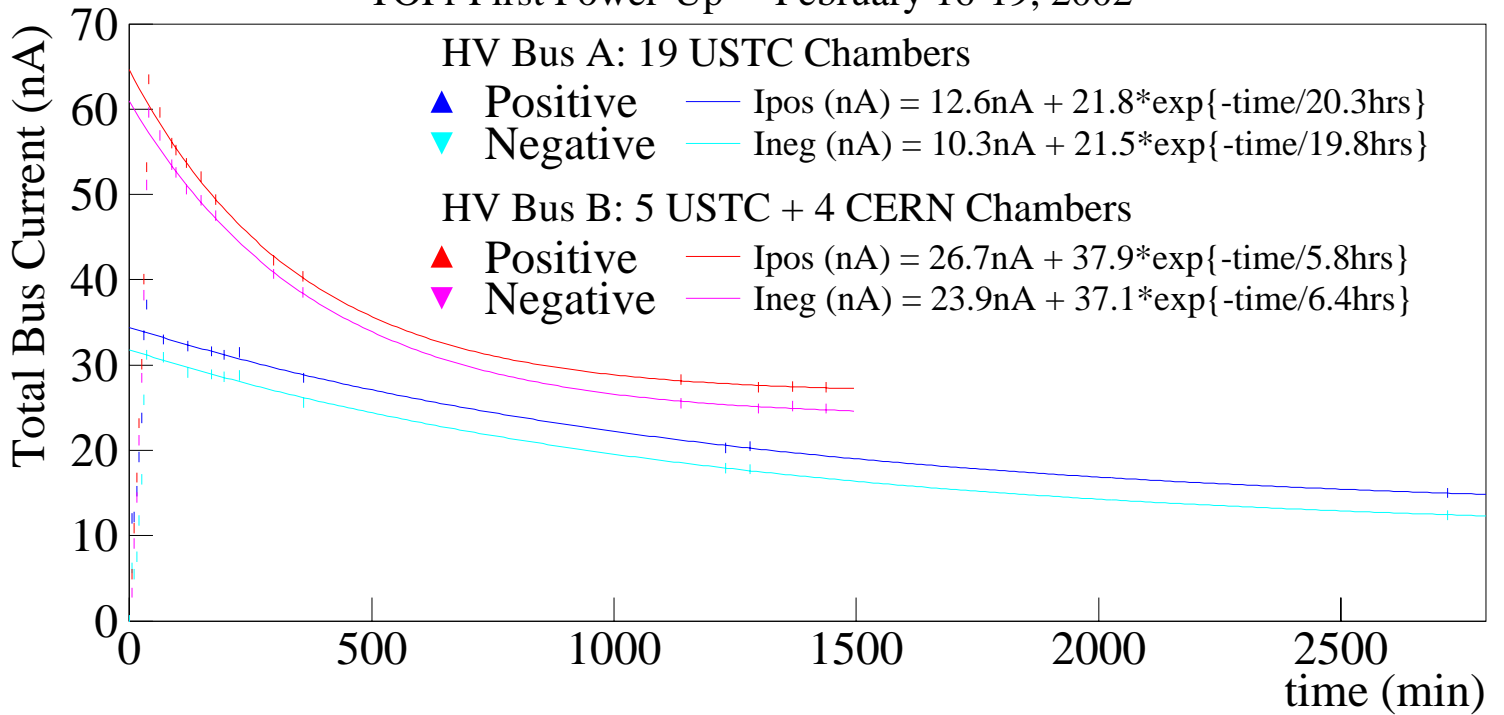
→ Realistic test of viability of this technology in an actual collider experiment.....





# TOFr tray appears to work as expected....

TOFr First Power-Up February 16-19, 2002



## Summary/Outlook:

TOF remains a viable technique for Particle Identification in modern experiments...

MGRPC detectors are dirt cheap and appear to outperform the conventional technology...

## Recent Major Successes

- a specific fishing line is a *great* choice for the 220  $\mu\text{m}$  spacer...
- Detector module design (Rice v.11) is now final for STAR... <60ps stop-resolution is typical...
- Maxim 3760 preamp & other standard components is an adequate approach to the FEE...
- Collaboration of US and CN institutions developed...
- A working prototype for STAR now exists!

## The basic long-range plan

Study performance of Single tray of MGRPC for STAR, “TOFr,” in test beams and in STAR...

Proposal for large-area system, “TOFR,” in Spring 2002...

Installation of TOFR in the 2004 shutdown...

## Hurdles Still Ahead of Us

- Solving the problems yet to be uncovered...
- Digitization for the large-area system must be “on-board”... (SBIR submitted 2/2002)  
TDC chip w/ required specs does not exist, although CERN efforts look promising...  
We have standing request for these when they become available...
- Marriage of preamp+comparator w/ the digitization on one board...  
Facing the age-old nightmare of analog and digital electronics on the same board...  
~6 EE-year engineering effort...
- Electronics R&D... (~500+ k\$?) spinoff applications exist...
- Building the large-area system itself.... (~5M\$ ? = 4M DOE + 1M CN-contributed)